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## The Multi-Channel Maximum-Likelihood (MCML) method: extension to multisource estimation and evaluation

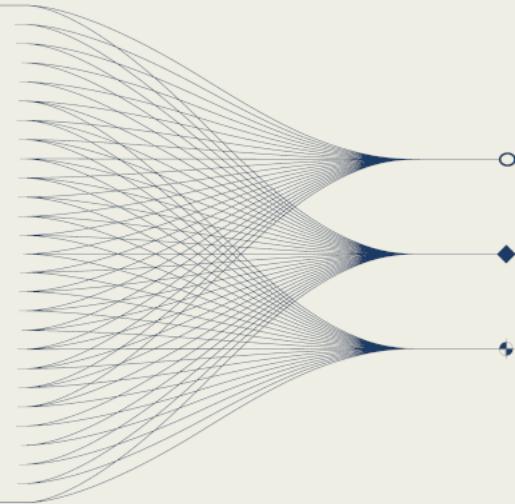
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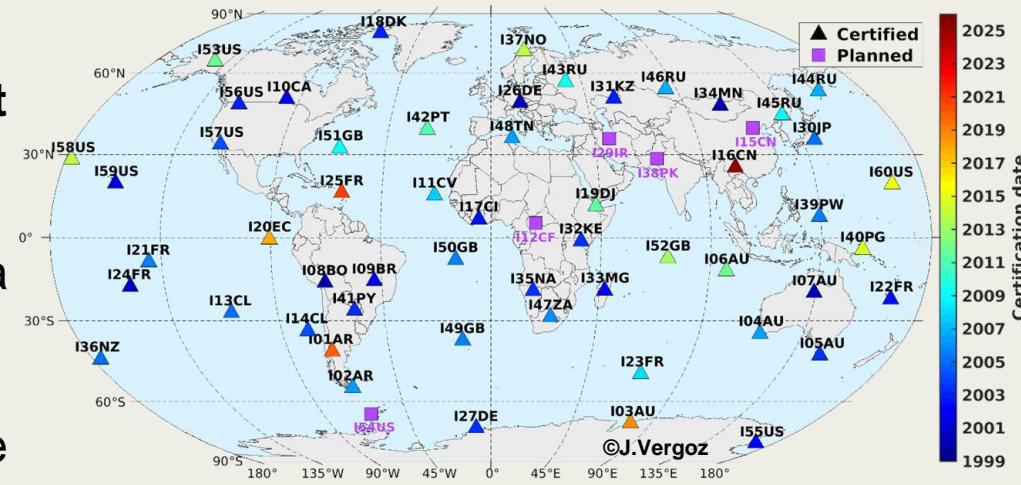


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## Introduction on the infrasound technology

- The international monitoring system (IMS) is designed to detect explosion of 1 kT eq. TNT everywhere at anytime
- Today, **90% (54/60)** of infrasound stations are **operational** on a total of 60 planned
- A wide range of natural and anthropogenic sources are detected
- Detecting explosions with low signal-to-noise ratio from diverse sources is a key challenge



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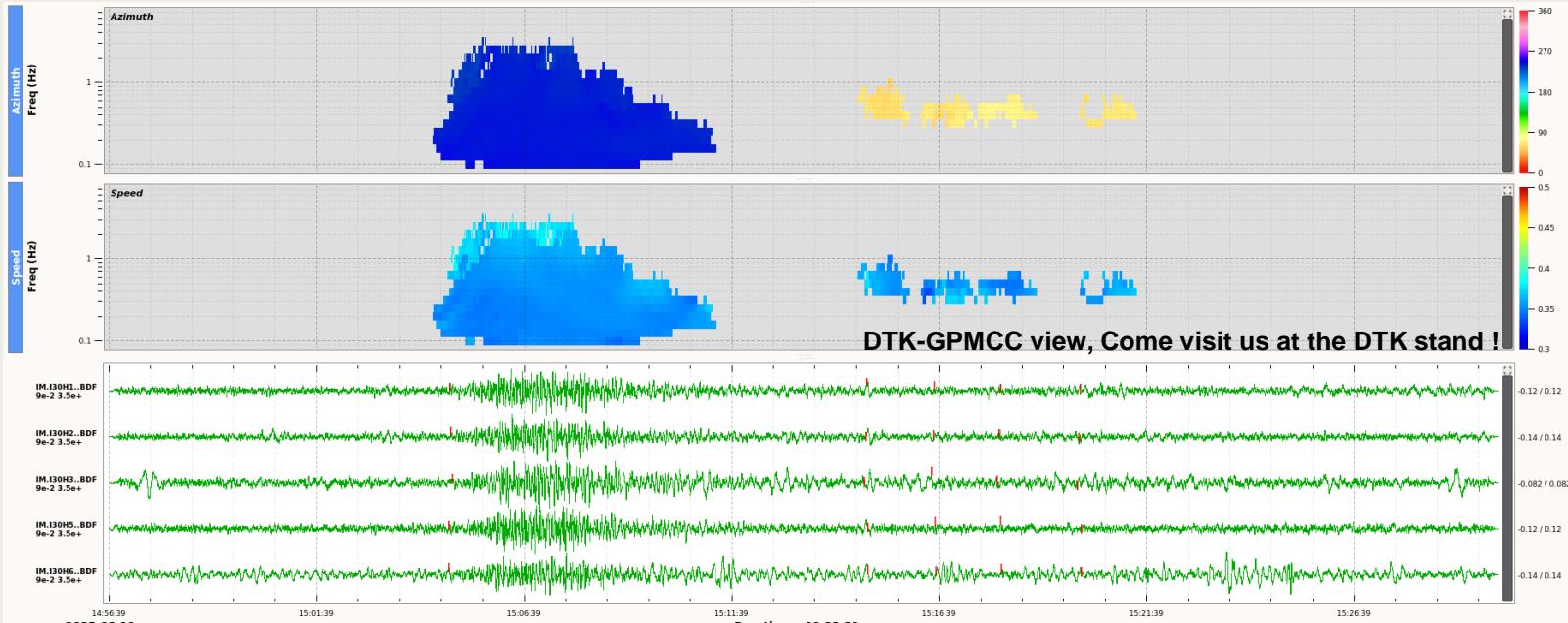
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## Overview of existing signal processing methods

Developments over time:

- F-detector, commonly used in NDCs (e.g. Arrowsmith et al., 2009)
- Principally used in a **single frequency band**
- Large variety of **signals covering a wide frequency band**

- Progressive Multi-Channel Correlation (Cansi, 1995; Garcés, 2013): detector (consistency) and estimator (TDOA)
  - Used **operationally since 2000 at the IDC** and distributed to our NDC partners
  - Feedbacks: loss of detection and estimation errors under unfavorable observation conditions



Fireball above Japan  
in 19/08/2025,  
detected at I30JP



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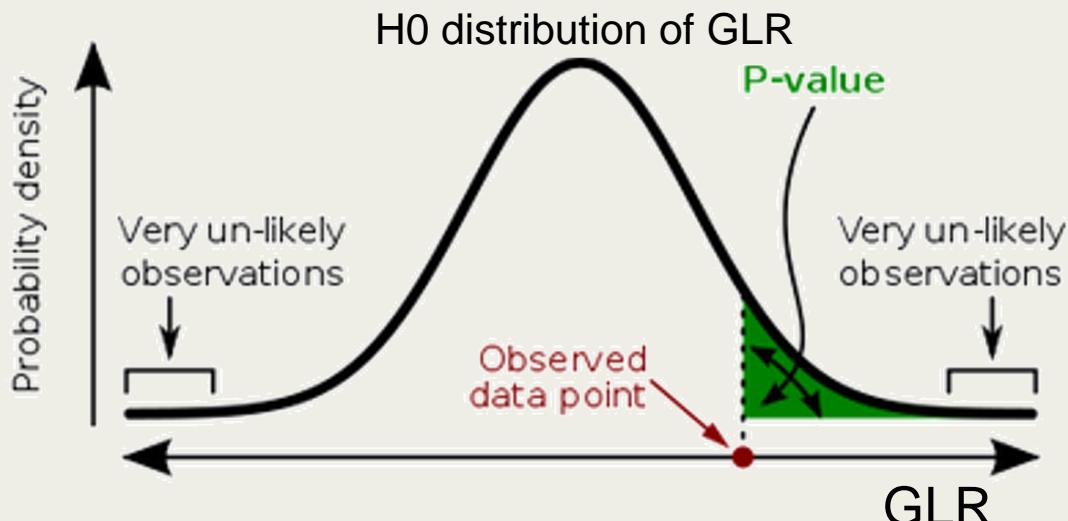
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## A new signal processing method: Multi-Channel Maximum Likelihood (MCML, Poste et. al (2023))

- Detection: Generalized Likelihood Ratio (GLR) (Burgess, 1994)
- Detection threshold by p-value, 1% of false positive
- Estimation: Maximization of the likelihood  $\mathcal{L}(\theta, \nu, s^2, \sigma^2)$  (MLE)
- Able to estimate the SNR and the uncertainty of the estimation

$$GLR = 2 \max_{\theta, \nu} (\mathcal{L}_* - \mathcal{L}_{0*})$$

$$p\text{-value} = \int_{GLR}^{+\infty} p_{GLR|H_0}(t) dt \in [0, 1]$$



$$R(\theta, \nu) = \mathcal{T} \Gamma_0 \mathcal{T}^\top \quad \text{with} \quad \mathcal{T} = \begin{bmatrix} -\nu \cos(\theta) & \nu \sin(\theta) \\ \nu^2 \sin(\theta) & \nu^2 \cos(\theta) \end{bmatrix}$$

$$\Gamma_0 = \frac{3(M \times SNR + 1)}{8\pi^2 \times MSNR^2 ((f_{max})^3 - (f_{min})^3)(t_{max} - t_{min})} (Z^\top \Pi_M^\perp Z)^{-1}$$

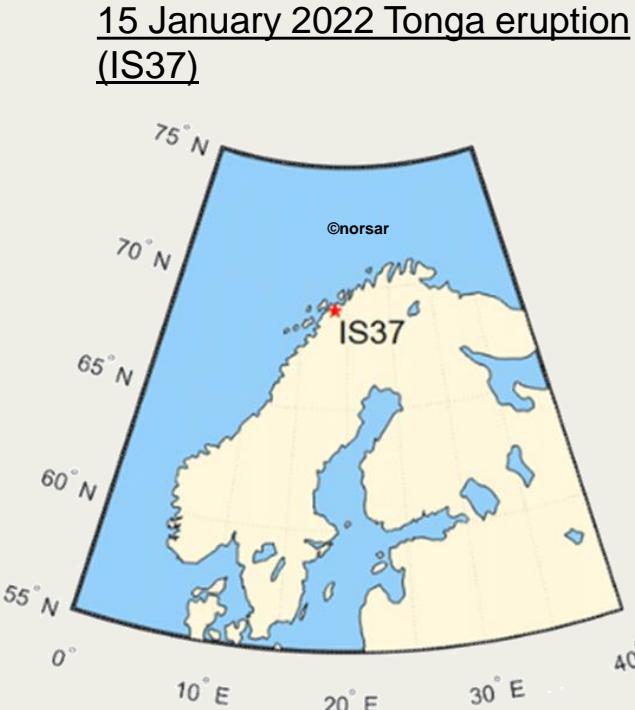
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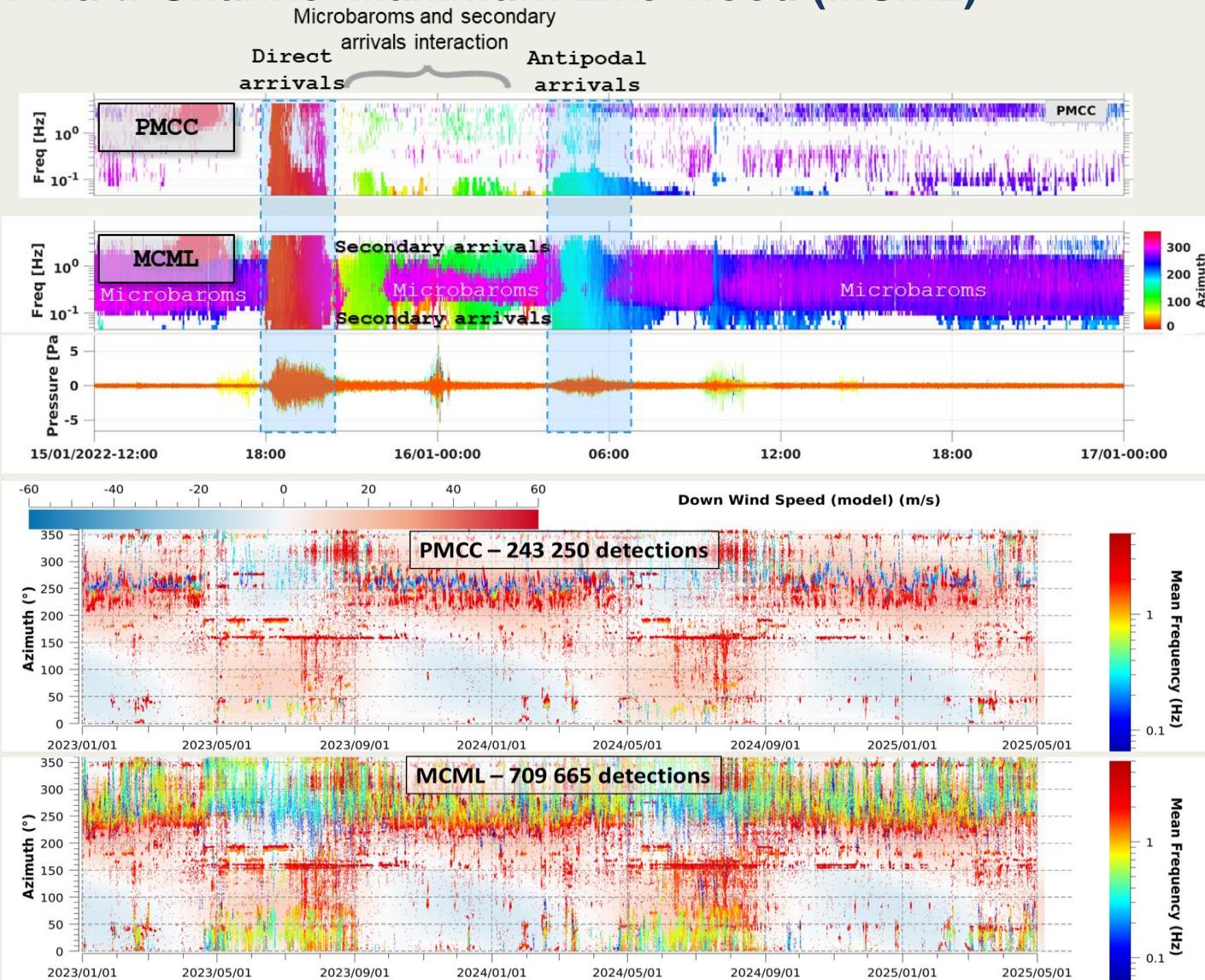
## A new signal processing method : Multi-Channel Maximum Likelihood (MCML)

**MCML enhances the detection of events of interest (Tonga arrivals) amid multiple coherent sources**

MCML enables more numerous and continuous detections revealing seasonal variations of stratospheric guides under low SNR conditions

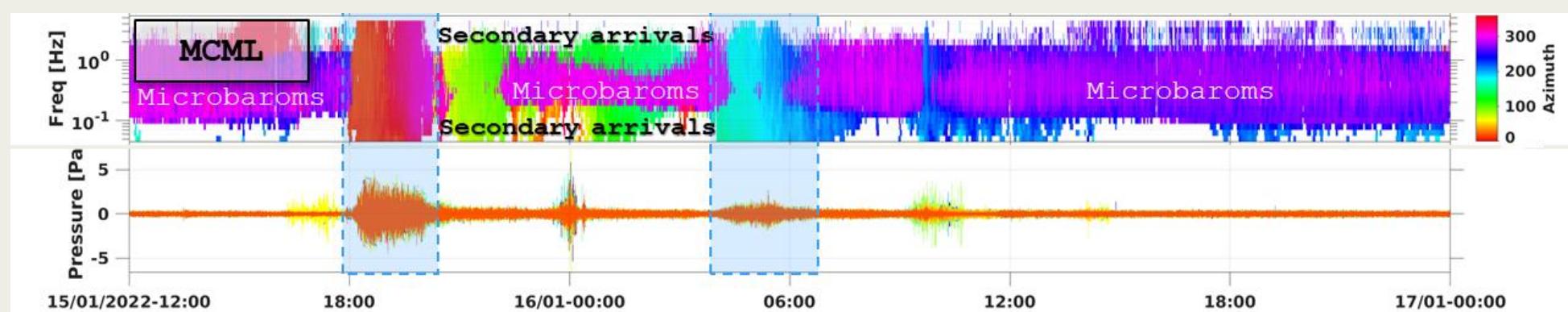


Signal processing from 2023 to 2025 at IS37



## Development of MCML-based method for multisource estimation

- MCML outperforms PMCC under **unfavorable recording conditions** (SOI interferences, low SNR)
  - MCML has been optimised for **real time processing** and implemented in the same time-frequency space as PMCC
  - **Single source hypothesis is not valid in reality**, MCML detects the **dominant source**
  - Several multisource algorithms exist (MUSIC, Schmidt 1986; CLEAN from: Wang et al 2004 or Den Ouden et al 2020), **need to be implemented for real time processing**
- Towards a cost effective method: develop and evaluate multisource algorithms to the state-of-the-art





## Development of MCML-based method for multisource estimation

The signal model includes N infrasound sources recorded at a station with M sensors

$$z_m(t) = \sum_{n=1}^N s_n(t - \tau_m(\theta_n, v_n)) + n_m(t) \quad , n_m(t) \text{ be the noise at the sensor } m.$$

### Multi peak selection method:

- Using a 2D peak finder (C.A.V. Aguilera, 2016) to **select maxima** on the **monosource** likelihood function

### Iterative source signal deflation method:

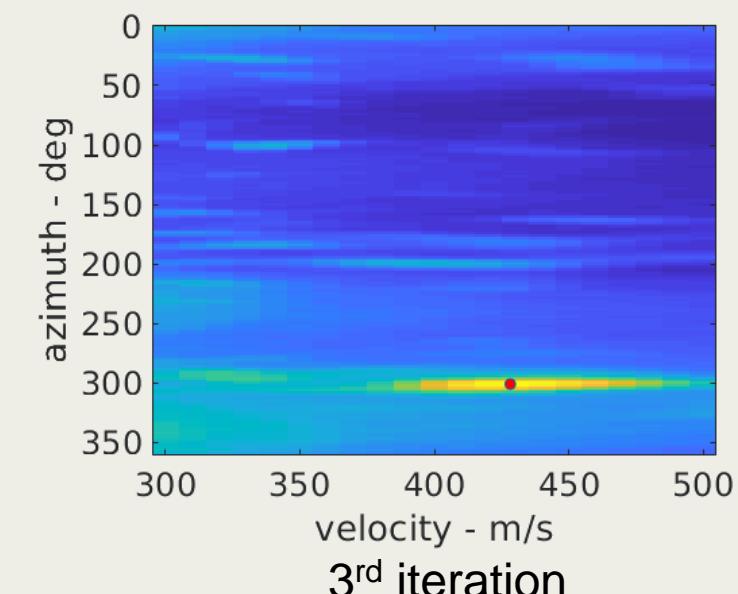
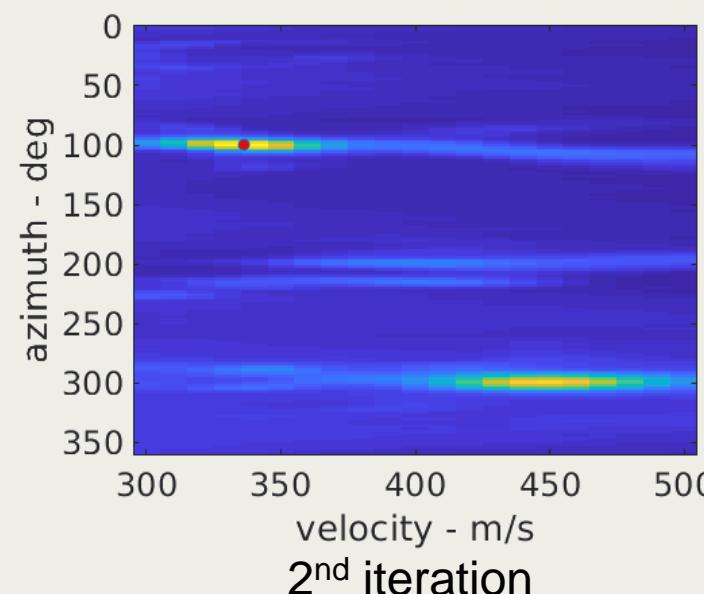
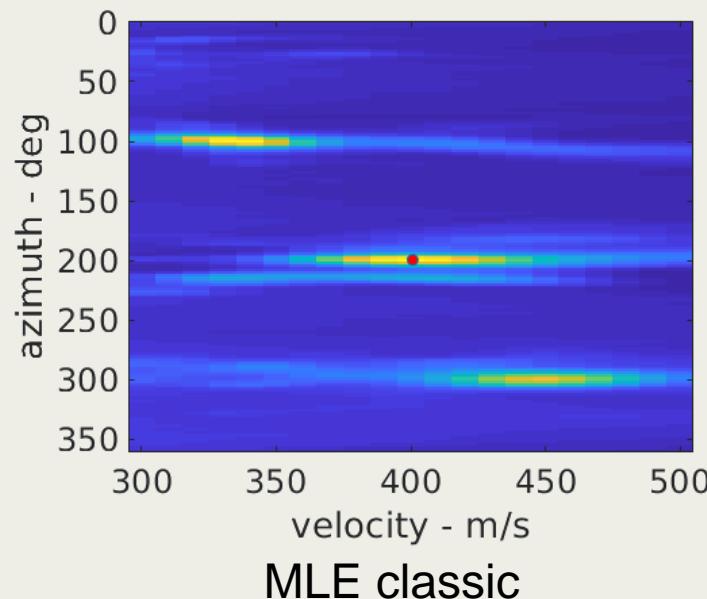
- Estimation of the first source and SOI:  $z_m\left(t + \tau_m(\widehat{\theta}_1, \widehat{v}_1)\right) = s_1(t) + n'_m(t)$



## Development of MCML-based method for multisource estimation

**Subtract a percentage** of the SOI (i.e 25%, Wang et al. 2004) and then iterate

View of the likelihood function on synthetic signal (red dot represents the maximum)





## Development of MCML-based method for multisource estimation

- The **spectral matrix model** of the monosource estimator is

$$C_1(f_i) = a(f_i, \theta_1, v_1)\sigma_{s1}^2 a^H(f_i, \theta_1, v_1) + \sigma_b^2 I_M$$

where  $\sigma_{s1}^2$  the power of the first source,  $\sigma_b^2$  the power of the noise,  $I_M$  the identity matrix

- Given the estimated parameters  $(\theta_1, v_1)$  from MCML, the spectral matrix is generalised as

$$C(f_i) = \sum_{n=1}^N \sigma_{sn}^2 a_n(f_i, \theta_n)a_n^H(f_i, \theta_n) + \sigma_b^2 I_M$$

- We use iteratively the **Whittle Likelihood as an estimator**

$$\mathcal{L}(\theta, v) = \sum_{f=f_{min}}^{f_{max}} \log \det C(f_i)^{-1} + Z^H(f_i)C(f_i)^{-1}Z(f_i)$$

where  $Z(f_i)$  is the DFT of the received signal

- Also implemented with approximations to **reduce computational cost** (theoretically calculated)

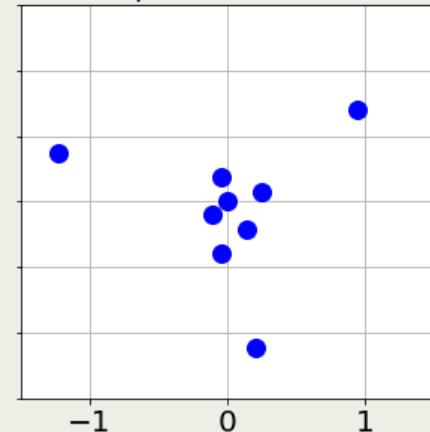


## Comparison and selection of a cost effective multisource estimator

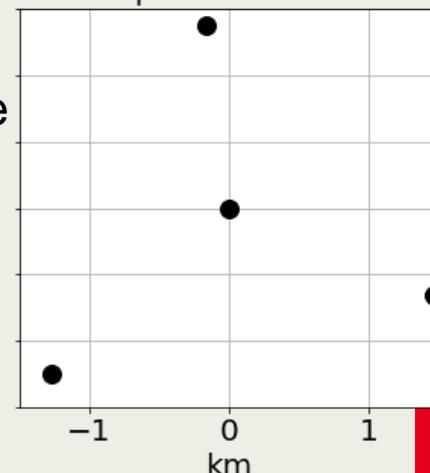
### Protocol:

- 3 sources in one signal with random azimuthal and velocity distribution
- Known number of sources
- Different frequency bands: [0.1 – 0.12] Hz, [0.5 – 0.65] Hz, [3.2 – 4] Hz
- Different geometries: IS31, IS46
- Different SNR between the sources: 5, 0, -5 dB
- A false estimation rate is determined by identifying the fraction of azimuth and trace velocity errors larger than the two-sigma theoretical uncertainty of MLE
- 1000 simulations for statistics significance

IS31: 9 sensors  
mean aperture = 1.43 km



IS46: 4 sensors  
mean aperture = 2.76 km



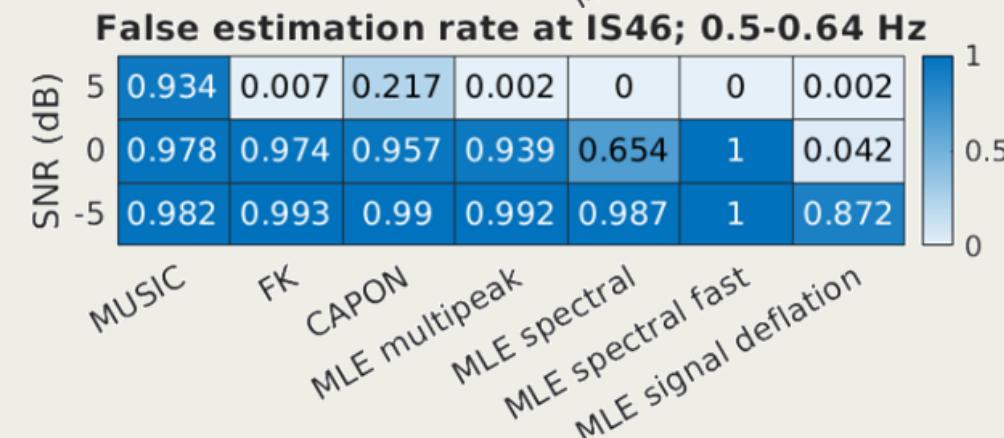
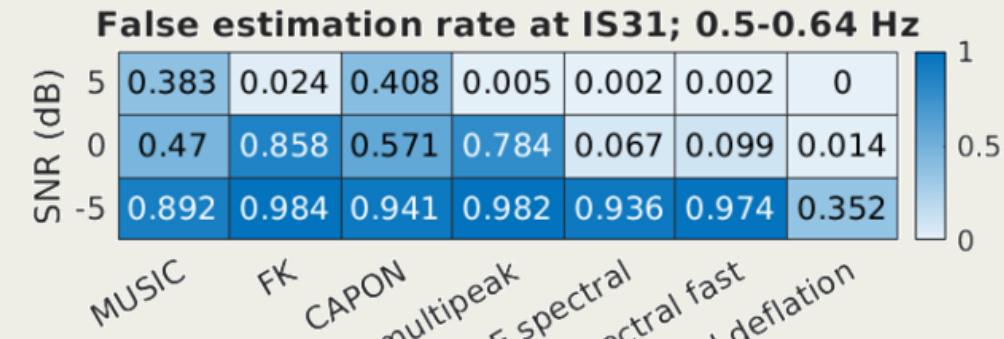


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## Comparison and selection of a cost effective multisource estimator

- False estimation rate in function of the SNR of sources and different algorithms for IS31 and IS46
- **MLE signal deflation with lowest false estimation rate**
- **Lower false estimation rate with increasing number of sensors**
- **MUSIC struggles with poor number of sensor**
- Other results: **Lower false estimation rate with increasing frequency**

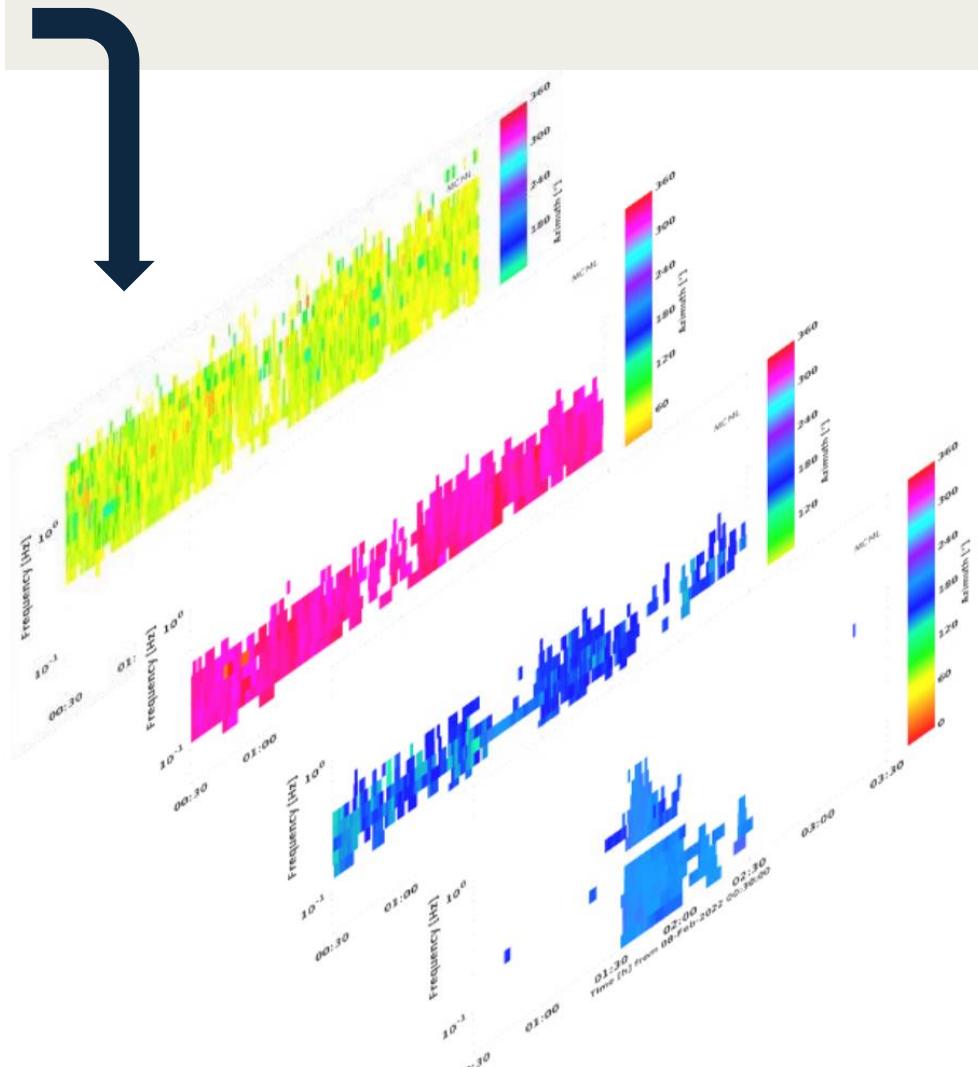
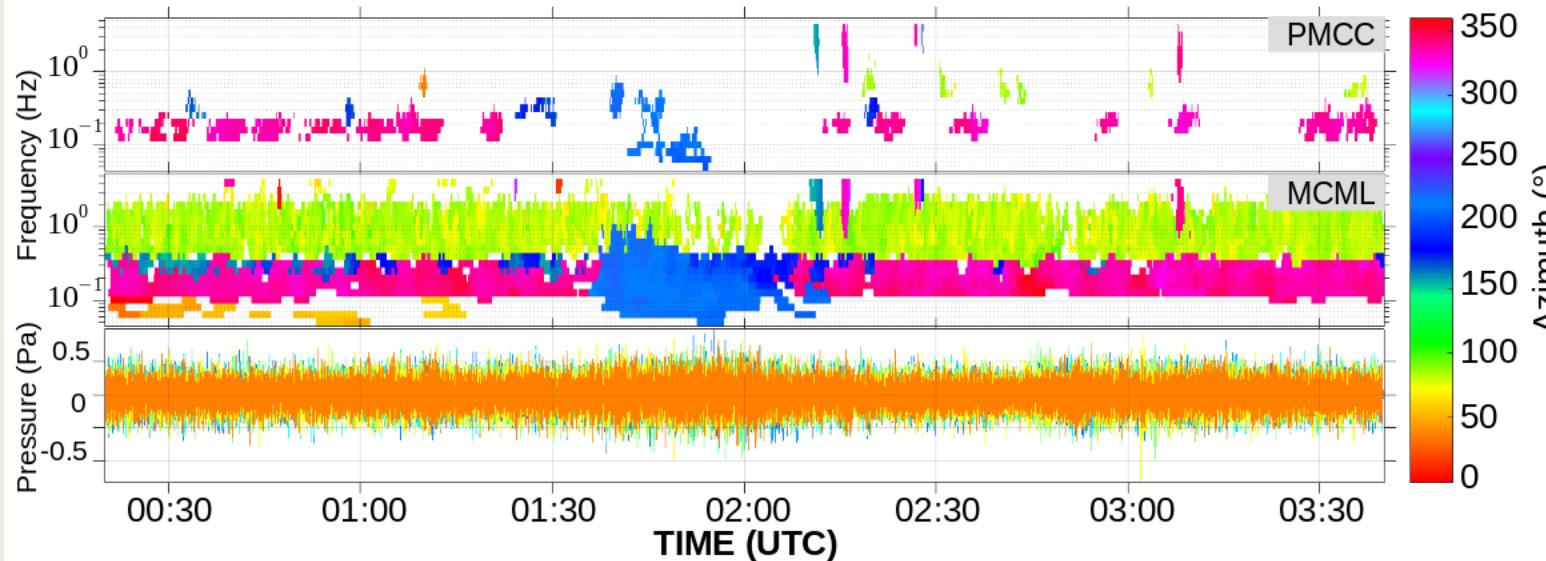




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## Comparison and selection of a cost effective multisource estimator



- Detection of the fragmentation of a meteor 500 km off the Namibia coast (February, 7th 2022) at IS19 (Djibouti)
- Removing **2 defective sensors** allows PMCC to detect
- Several sources interfering **detected continuously**: microbaroms (Atlantic & Indian now detected), surf noise and the meteor



## Conclusion

- **MCML outperforms PMCC**, especially at **low SNR** and amid **multiple interfering SOIs**
- Currently undergoing operational testing at the French NDC
- MCML is slated for **implementation in DTK tools** in 2026-2027 becoming the new operational detector
- This improvement leads to the development of multiple sources estimators
- **MCML with iterative signal deflation** offers a cost-effective solution for **real time processing**
- Optimise the proportion  $\alpha$  of subtracted signal in MCML multisource
- More **systematic evaluations are needed** on reference events
- **Need of a robust post-processing method** (pixel clustering) to build event



# Thanks for listening !

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