

Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

TMc Student / University of Brasília / Seismological Observatory



Presentation Date: 11 September 2025



Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

02.2-767

## **Outline**

- 1. Introduction
  - 1.2 Scientific Context
- 2. Methodology
- 3. Data Analysis and Results
  - 3.1 Event of July 28, 2022
  - 3.2 Event of December 6, 2019
- 4. Discussion and Conclusions





Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

02.2-767

## Introduction

#### IMS Infrasound Stations

- Operated by CTBTO, designed for nuclear test monitoring.
- Also detect natural and anthropogenic events (e.g., bolides).

## **Undergraduate Research**

- Analysis of bolides detected by IMS stations in South America (2018–2025).
- Data source: NASA CNEOS catalog.

#### Case Studies

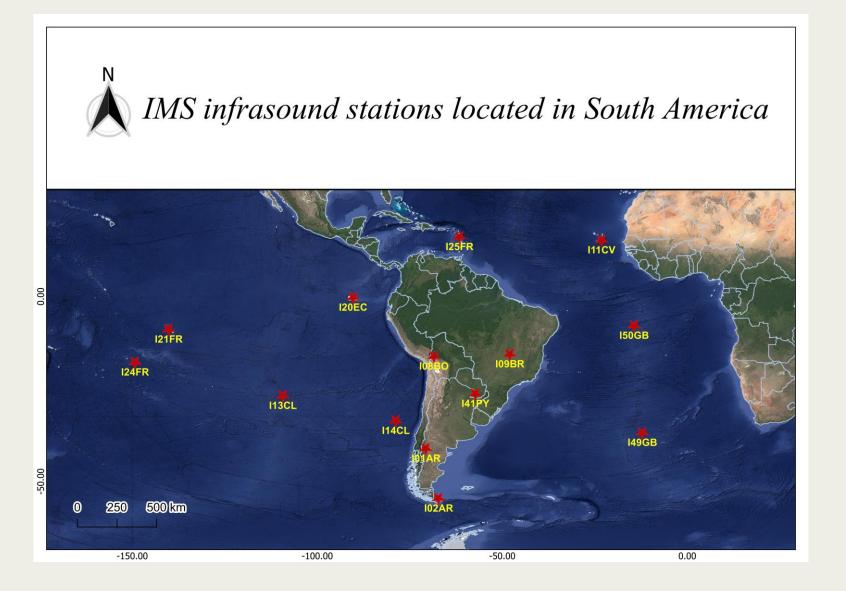
- Two events: 2019 and 2022.
- Energy < 1 kt TNT.





Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767





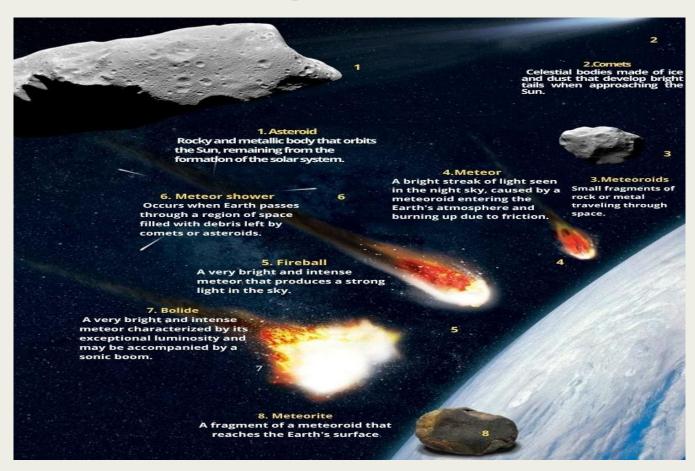


Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

02.2-767

# How is infrasound generated?

# Scientific Context



Modified from the American Meteor Society (American Meteor Society, 2025).

## Meteoroids in the Atmosphere

- Fragments of asteroids or comets.
- Produce visible phenomena: meteors, fireballs, bólides.
- Classification depends on mass, velocity, and energy.

#### **Infrasound Production**

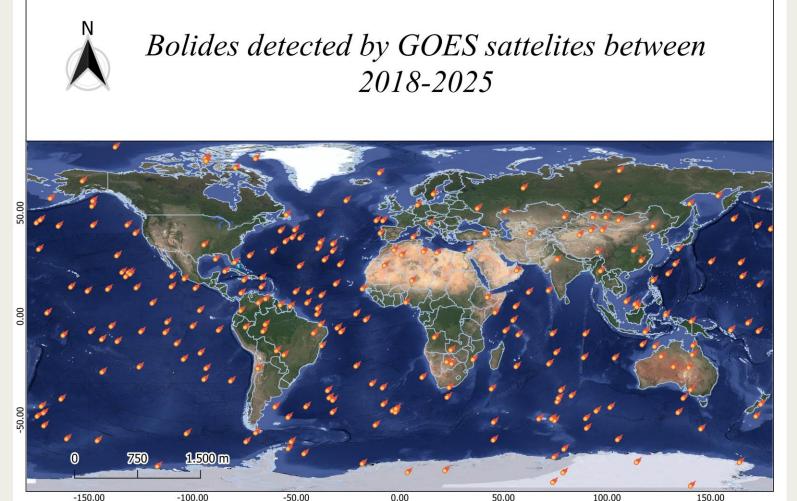
- Atmospheric entry generates shock waves.
- Shock waves propagate as infrasound (<20 Hz).



Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

# Methodology



- •Data Source: CNEOS database
- Detection System: GLM sensors onboard GOES satellites.
- •*Time Period*: 2018 2025
- •Data Collected: 263 bolide events recorded worldwide.
- Purpose: Analyze global distribution and characteristics of bolide events.



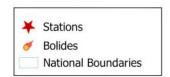
Letícia Guedes Assunção , Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

# Methodology



# Bolides detected by IMS infrasound stations located in South America



- -150.00-100.00
- 263 bolides (2018–2025) recorded globally.
- by IMS infrasound • 19 detected stations in South America, 14 in SEL-3 Bulletin and 5 via waveform analysis.
- Case Studies: Dec 6, 2019 (Brazil) & Jul 28, 2022 (Pacific).
- IMS Infrasound Stations: 120EC (Ecuador), I08BO (Bolivia), I09BR (Brazil), and I41PY (Paraguay).
- Analysis with DTK-GPMCC: direction, speed, frequency and signal duration.





Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

# Data Analysis and Results

# Event of December 6, 2019



# Bolide Event - Fortaleza, Brazil

- Detection time: 10:28 UTC (CNEOS).
- Altitude: ~19.5 km (CNEOS).
- Energy released: ~0.15 kt TNT (CNEOS).

#### IMS infrasound detections:

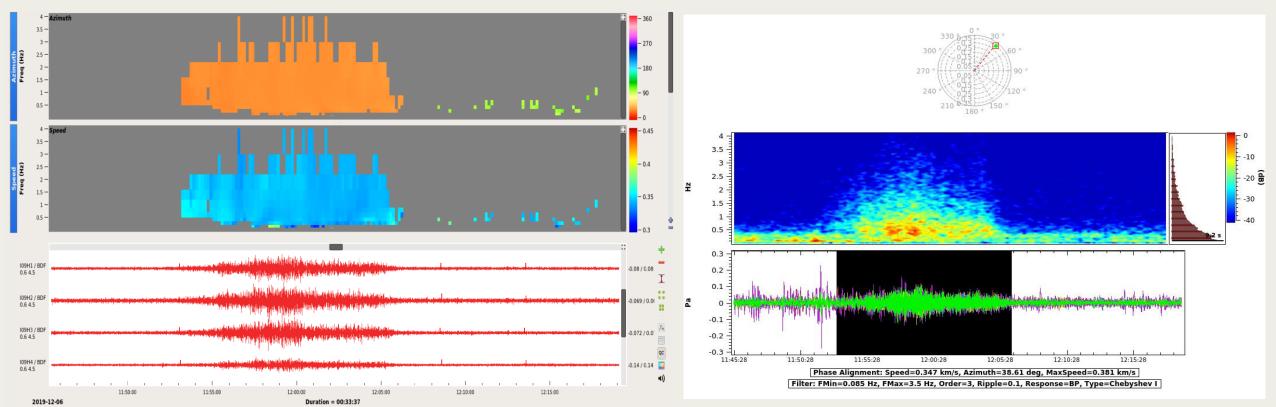
- I09BR (Brazil)
- I41PY (Paraguay)
- I08BO (Bolivia).



Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

02.2-767

# 109BR Station



**I09BR** → closest station, ~1,800 km.

- •Waveform → arrival at 1h27min, duration ~13 minutes.
- •Spectrogram → energy 0.3–1 Hz, peak at 0.5 Hz (stratospheric range).
- •Back-azimuth (39°) → source Direction.
- •Apparent velocity (0.346 km/s) → confirms stratospheric propagation.

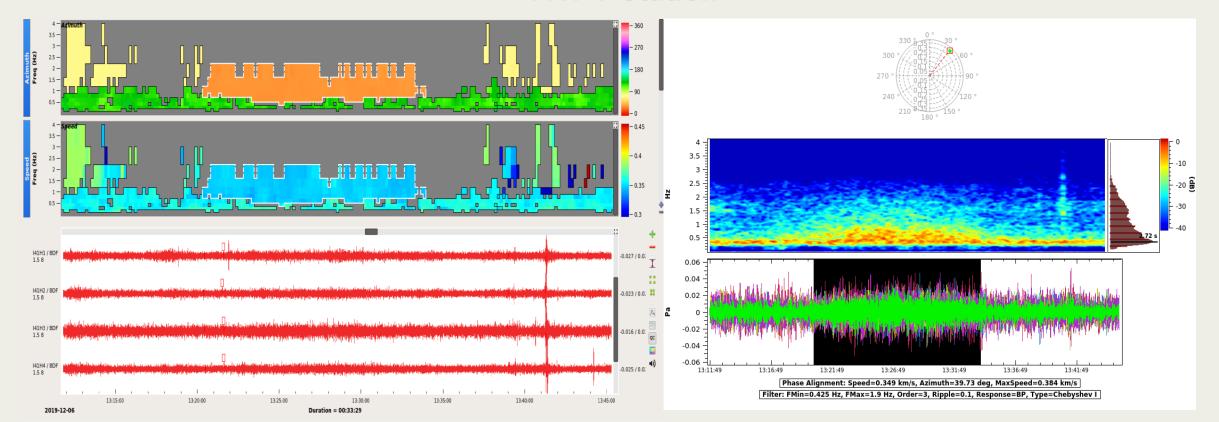




Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

## **I41PY Station**



**I41PY** → second station, ~3,300 km

- •Waveform → arrival at 2h41min, duration ~14 minutes
- •Spectrogram → energy 0.3–1 Hz, peak at 0.7 Hz (stratospheric range)
- •Back-azimuth (39°) → source direction
- •Apparent velocity (0.349 km/s) → consistent with stratospheric propagation

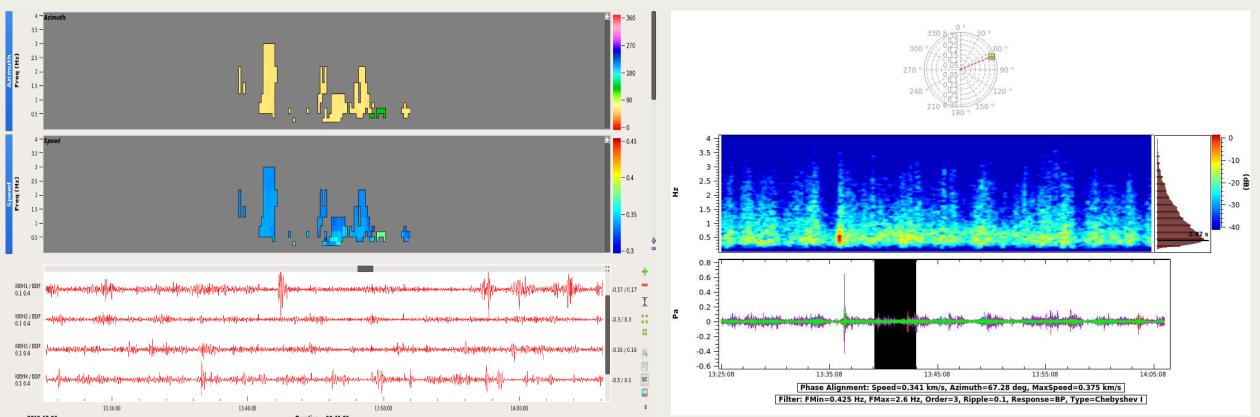




Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

# 108BO Station



**I08BO** → most distant station, ~3,600 km

- •Waveform → arrival at 2h57min, duration ~6 minutes.
- •Spectrogram → weak signal, energy around 0.5 Hz.
- •Back-azimuth (68°) → source Direction.
- •Apparent velocity (0.341 km/s) → consistent with stratospheric propagation.

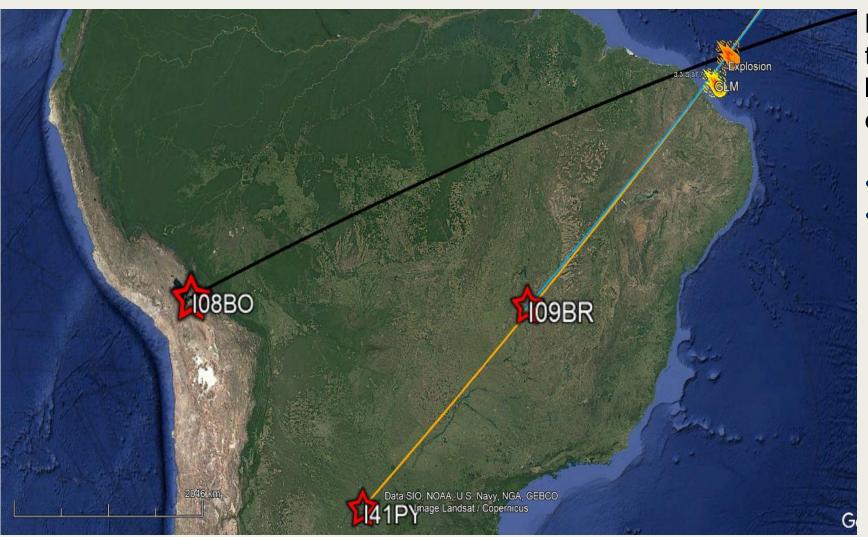




Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

# Localization



Based on the back-azimuths from the three IMS stations, the location of the atmospheric explosion was estimated as:

·Latitude: 1.6° S

Longitude: 36.8° W



Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

# Data Analysis and Results

# **Event of July 28, 2022**

Map Elements



# **Bolide Event**: Southeast of the Galápagos Islands.

- Detection time: 01:36 UTC.
- Altitude: ~37 km.
- Velocity: ~30 km/s.
- Energy released: ~0.68 kt TNT.

#### IMS infrasound detections:

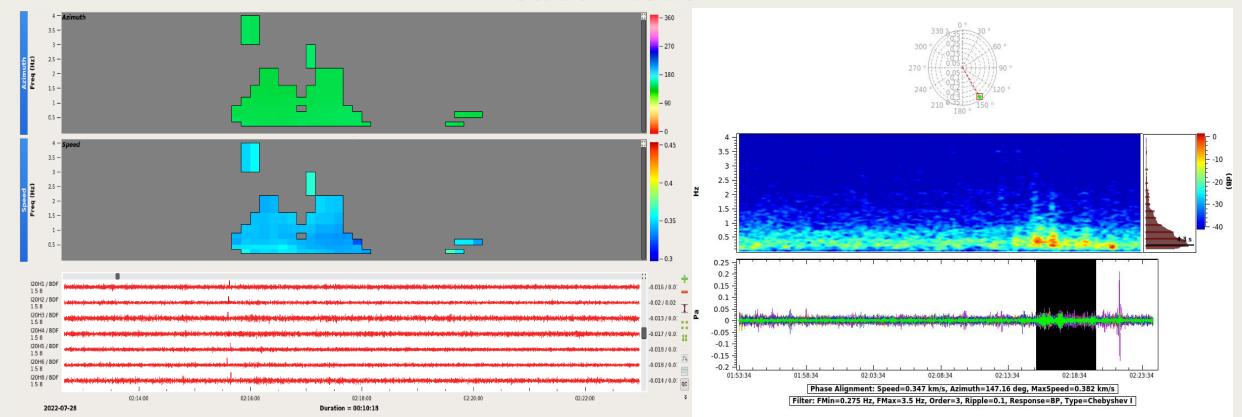
- Ecuador, Bolivia and Brazil.
- Signal tracked up to: 4,000 km



Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

02.2-767

## Station I20EC



**I20EC** → closest station, ~670 km

- •Waveform → arrival 39 minutes after GLM, duration ~4.7 minutes
- •Spectrogram → energy 0.3–0.5 Hz, peak at 0.5 Hz
- •Apparent velocity → 347 m/s
- •Propagation → tropospheric (lower atmosphere)

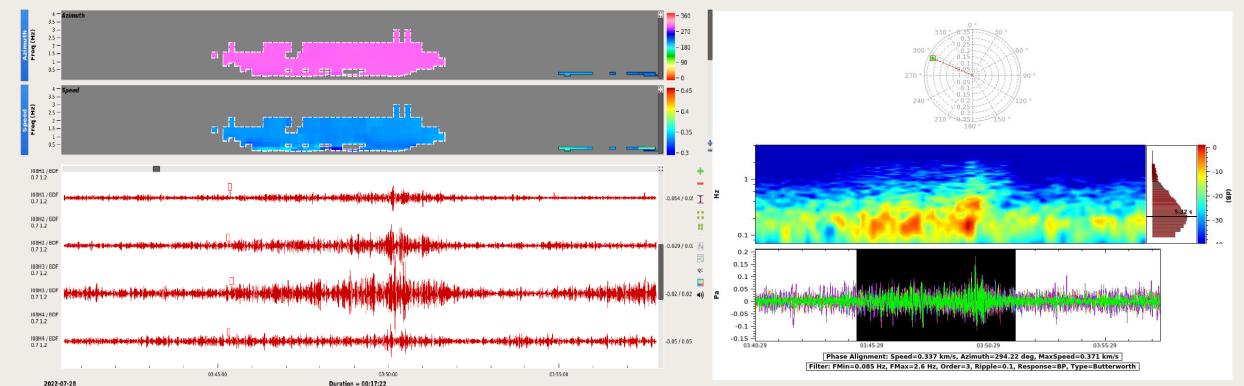




Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

## Station 108BO



**I08BO** → second closest station, ~2,378 km.

- •Waveform → arrival ~2 hours after bolide, duration ~7 minutes.
- •Spectrogram → energy 0.1–0.6 Hz, dominant 0.4 Hz.
- •Apparent velocity → 0.337 km/s.
- •Back-azimuth → 294°.
- •Propagation → stratospheric (upper atmosphere, long distance).

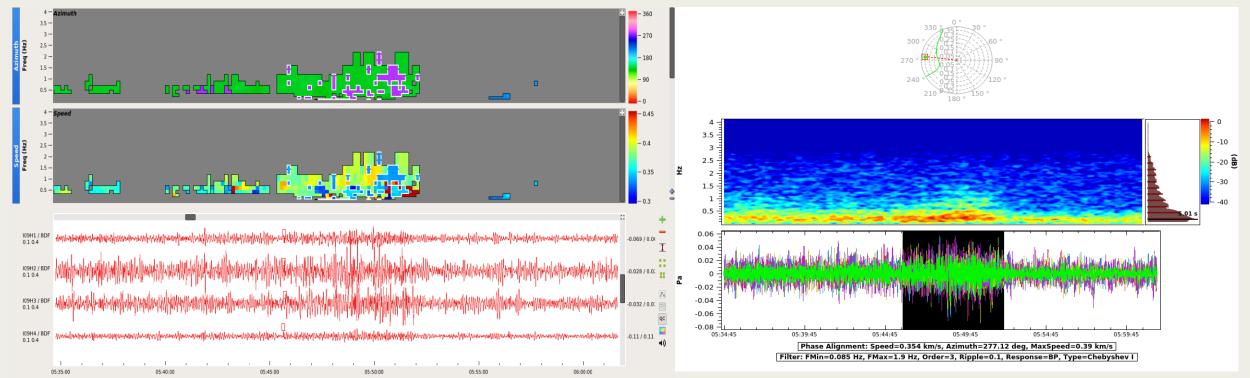




Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

## Station 109BR



**I09BR** → most distant station, ~4,450 km

- •Waveform → two signal families:
  - 277°, duration ~6.5 minutes, apparent velocity 0.354 km/s (matches true azimuth); 137° → not consistent with bolide
- •Spectrogram → energy at 0.56 Hz, band 0.2–0.6 Hz
- •Propagation → stratospheric (over 4,000 km).

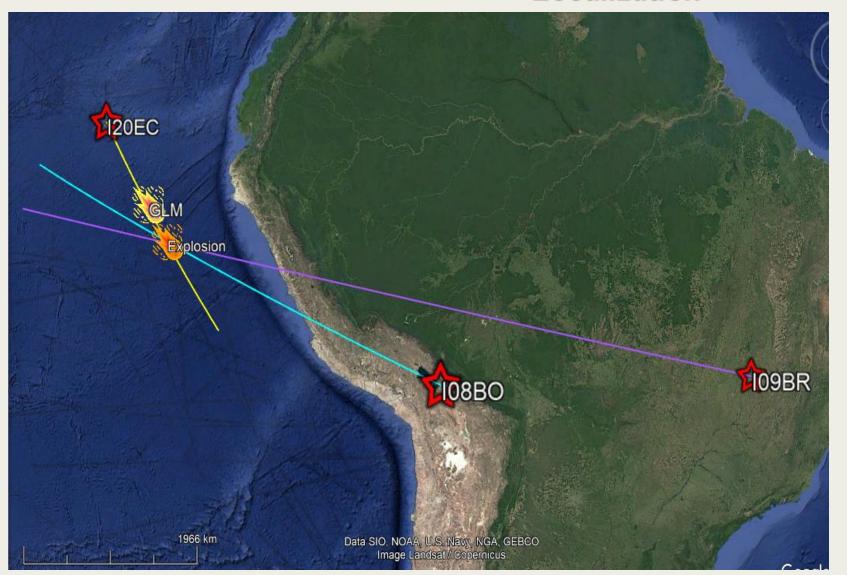




Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

O2.2-767

# Localization



The estimated coordinates are:

Latitude: 8.1° S

Longitude: 85.4° W







Letícia Guedes Assunção, Lucas Vieira Barros, Darlan Portela Fontenele, Nicolas Rainner Imanna, Juraci Mario de Carvalho, Ana Clara De Araujo Sousa

02.2-767

#### Discussion and Conclusions

# **Bolide Event Analyses – Dec 2019 & Jul 2022**

- IMS infrasound stations detect low-energy meteoroids over vast distances. Propagation:
- 2019 → mostly stratospheric.
- 2022 → tropospheric + stratospheric paths.
- Consistent with GLM sensor data  $\rightarrow$  validates IMS detection method.
- Demonstrates IMS potential for global meteoroid monitoring.



