

# On the use of metamodels for Bayesian localization and characterization of infrasound events

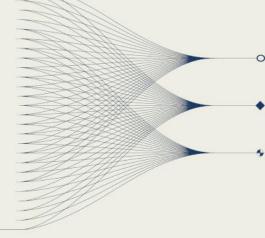
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#### ·························INTRODUCTION AND MAIN RESULTS

This work investigates the use of generalized Polynomial Chaos (gPC)-based metamodels, embedded in a Bayesian framework, to achieve more robust and cost-efficient source localization. Application to the 2018 Bering Sea bolide event: (1) reduced localization error from ~220 km (IDC estimate) to less than 50 km, consistent with NASA satellite observations; (2) highlighted the critical impact of arrival-time extraction uncertainty, comparable to atmospheric uncertainties; (3) revealed the ambivalent role of station IS59, whose contribution improves localization but also introduces significant uncertainty, illustrating the need of adaptive station weighting.



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### From forward modelling to inference

Source localization remains challenging due to atmospheric variability, uncertainties in signal picking, and the computational cost of full-wave propagation models. Bayesian inference provides a principled framework for combining detections with prior knowledge (here provided by IDC), but its practical use requires fast and accurate forward models.

Our goal: leverage Polynomial Chaos Expansions (PCE) as an efficient surrogate model for Bayesian localization. The arrival time at station  $\#_j$  is emulated by:

$$t_i(\mathbf{x}) \simeq \sum_{k=1}^K a_{ik} H_k(\mathbf{x}),$$

 ${\bf x}$  denotes the source location,  $\{H_k\}$  is the PCE basis, and  $\{a_{jk}\}$  is the set of coefficients computed using least square regression from simulated observations

$$\mathbf{a}_{j} = \operatorname{argmin}_{\mathbf{a}_{j} \in \mathbb{R}^{p}} \left\| \mathbf{t}_{j} - \mathbf{H} \mathbf{a}_{j} \right\|_{2},$$

where  $\mathbf{a}_j = [a_{jk}]$ ;  $\mathbf{H} = [H_k(\mathbf{x}_m)]$  and  $\mathbf{t}_j = [t_j(\mathbf{x}_m)]$  are computed from a sampling of source locations.

Assuming that arrival time measurement errors at distinct stations **are independent**, the joint likelihood  $P(t|\mathbf{x})$  factorizes so as to give

$$\prod_{j} P(t_{j}|\mathbf{x}) \propto \prod_{j} \sigma_{j}^{-1} \exp[-\epsilon_{j}^{2}(\mathbf{x})/(2\sigma_{j}^{2})],$$

where  $\epsilon_j$  is the difference between the observed arrival time and the PCE prediction.

## Application to the Bering sea bolide

On Dec. 18, 2018, a small asteroid entered Earth's atmosphere over the Bering Sea. We apply the Bayesian method using 6 IMS stations to test its ability to reproduce the reference locations (NASA [1]).

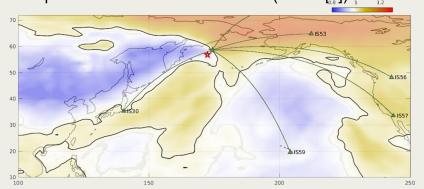
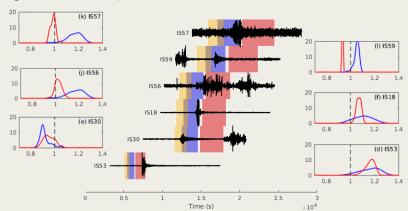


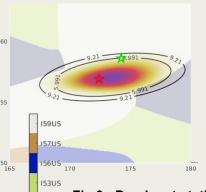
Fig. 1 : Eff. sound speed ratio; source locations, IDC: ★, NASA: ★.



**Fig. 2 :** Signals; **W**, **S**, and **T** phases using an empirical model [2]. Eff. sound speed ratios: source-level and path-averaged profiles [3].

# On the role of each station in the localization: likelihoods & posteriors

Each IMS station contributes differently to the likelihood surface. By examining partial likelihoods  $P(t_j|\mathbf{x})$ , we identify which stations **dominate the localization**.



Dominance of IS59 and its limitations: posterior map shows that IS59 apparently dominates the localization, but using the Eff. Sound speed profile at source level (see Fig. 2), which overestimates the efficiency of ducting.

**Fig.3 : Dominant station** -  $t_j$ 's are extracted from waveforms as the times for which the amplitudes of the rms first exceed a fixed threshold of 30%.

### **Future work**

Accurate uncertainty estimation using MCMC will target both arrival time extraction and the impact of background atmospheric variability. The MCMC scheme should **adaptively downweight** stations whose contributions are uninformative.

- [1] https://cneos.jpl.nasa.gov/reballs/
- [2] Blom et al., Improved Bayesian Infrasonic Source Localization for regional infrasoundGJI, 2015.
- [3] A. Goupy, A Spectral Approach of Multi-Scale Metamodelling Applied to Acoustic Propagation, PhD these, ENS Paris-Saclay, 2021.

