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QUANTIFYING BOLIDE ENERGY THROUGH INFRASOUND ANALYSIS: A CASE STUDY OF THE 2023 AUSTRALIAN EVENT

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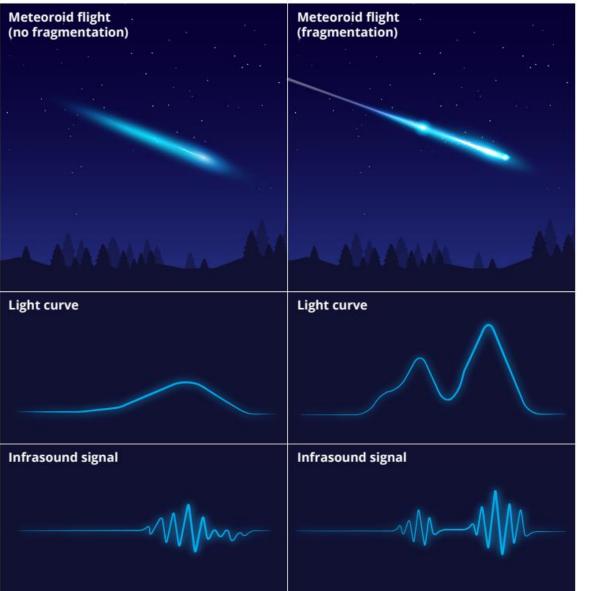


WHERE INNOVATION BEGINS

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OBSERVATIONS OF LARGE BOLIDES



Ongoing effort to better characterize bolides

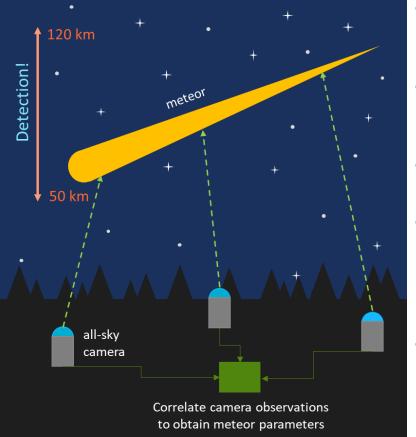
Events recorded and documented through various means of observation

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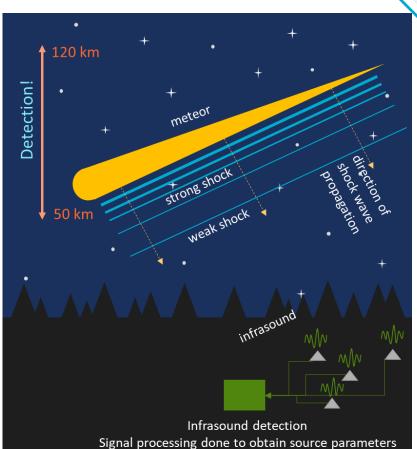
- US Government sensors
- GLM
- All-sky cameras (still and video)
- Radar
- Casual witnesses
- Other less conventional methods (infrasound & seismic)
- There is no perfect approach that would provide all answers we seek
- Many questions remain unanswered, and therefore it is imperative to leverage all approaches

Figure adapted from Silber (2024)

MULTI-MODAL DETECTIONS: ALL-SKY CAMERAS AND INFRASOUND

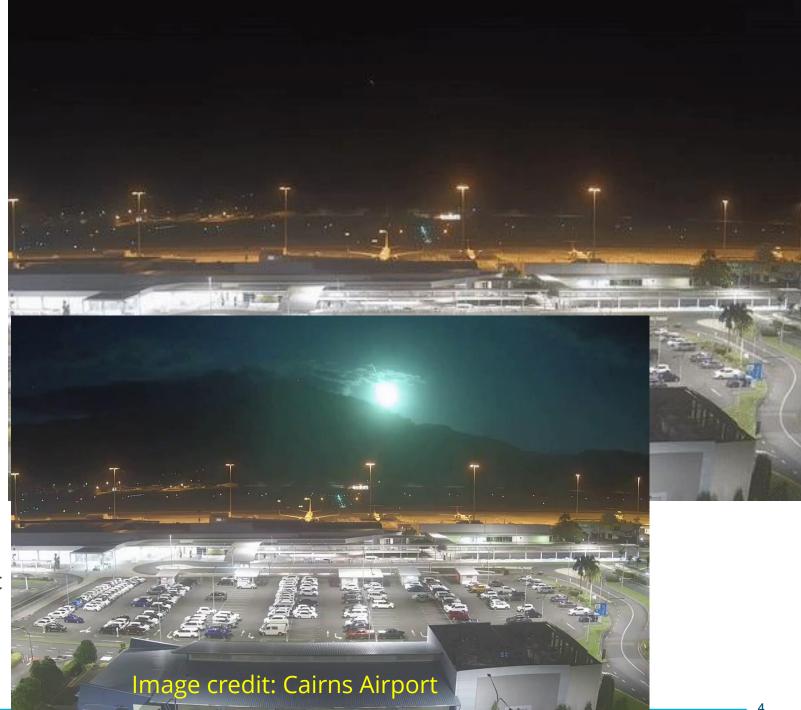


- Optical observations from multiple vantage points can provide valuable ground truth information.
- Accurate trajectory, velocity, entry angle, etc. can be fed into infrasound search algorithm (Silber 2024).
- Infrasound detection provides additional layer of information about the source.
 - Detailed studies using multi-modal detections of fireballs can provide information (and model validation and refinement) that can be later used for events with limited ground truth.
 - The Australian bolide: the Desert Fireball Network of cameras, a project led by Curtin University, observes meteor activity in Australia



AUSTRALIAN BOLIDE (MAY 20, 2023)

- Bolide over Queensland, at 9:22 pm local time (May 20, 2023)
- Was visible 600 km away
- Blackbull, a small rural locality between the Gulf communities of Normanton and Croydon, in northwest Queensland
- Many people captured the moment the green-blue fireball entered the atmosphere until the object disintegrated in a flash of orangeyellow light
- Queensland bolide is the largest event over Australia since 1988

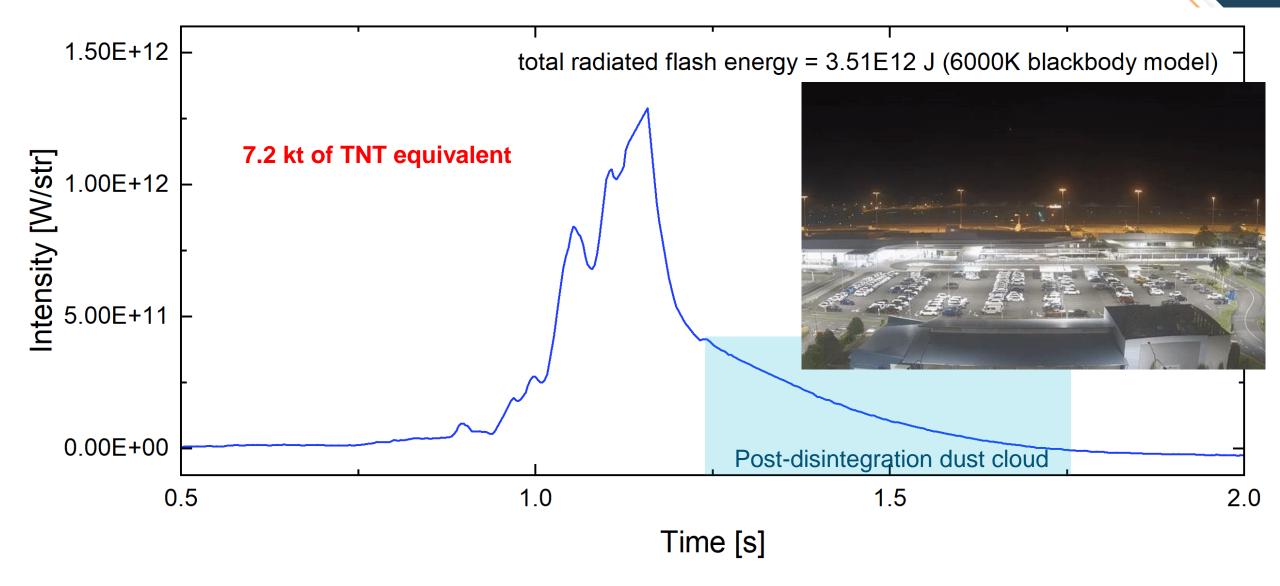


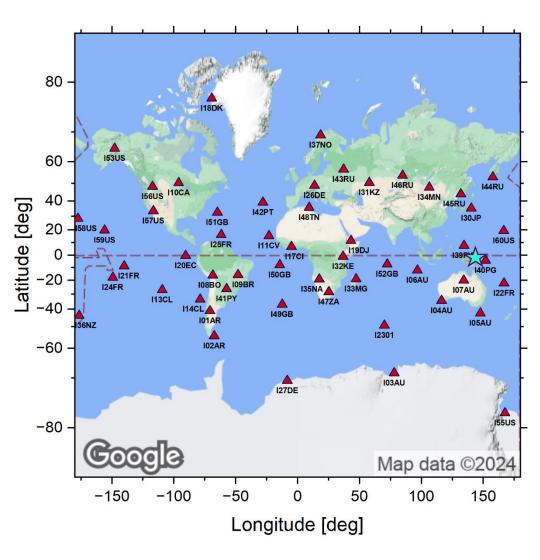
AUSTRALIAN BOLIDE (MAY 20, 2023)

- Altitude = 29 km, v = 28 km/s
- Entry angle = 39 degrees
- Size ~ 3.5 m
- Produced a sonic boom
- E = 7.2 kt of TNT equivalent
- Top 20 in the entire JPL CNEOS database, and the most energetic bolide over Australia since 1988 (since US government started reporting bolide events)



USG SENSOR OBSERVATIONS AND LIGHT CURVE

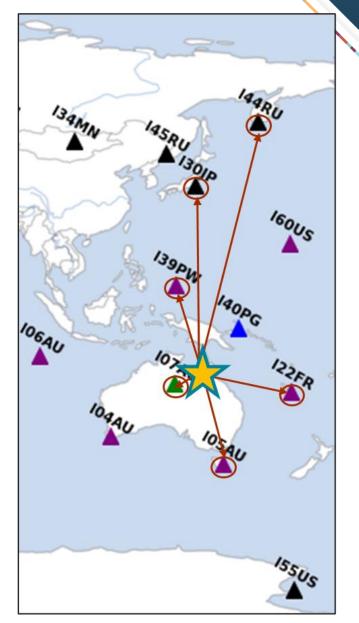




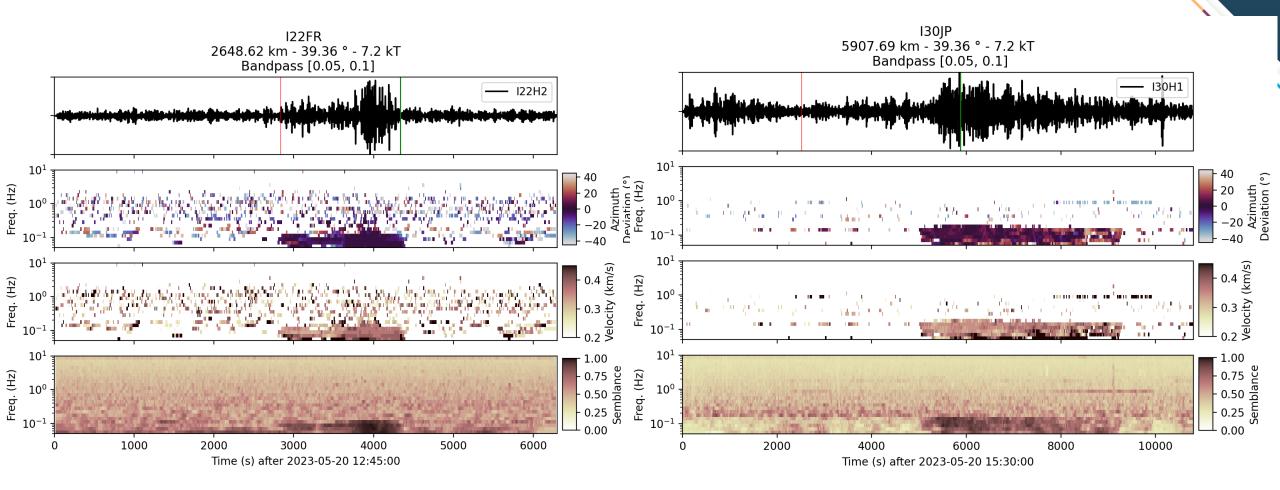
INFRASOUND DETECTIONS

- I07AU 830 km (weak detection)
- I22FR 2650 km
- I05AU 2800 km
- I39PW 2930 km
- I30JP 5907 km
- I44RU 8030 km

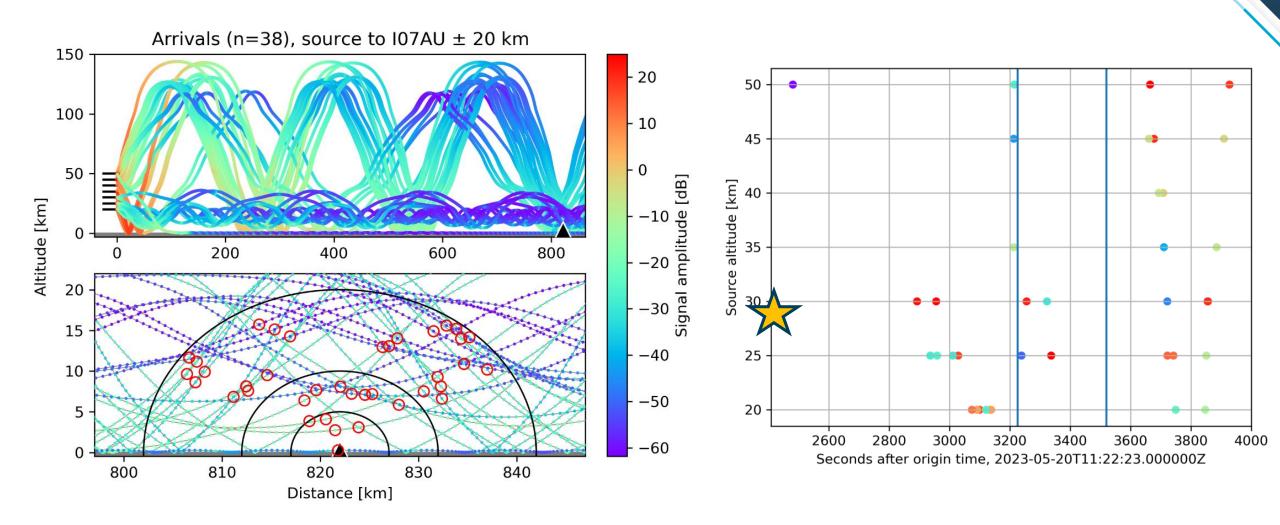
- Searched for infrasound on IMS stations of the CTBTO PrepCom
- Signals detected on 6 stations of the IMS network, predominantly east and north relative to the bolide



INFRASOUND DETECTIONS

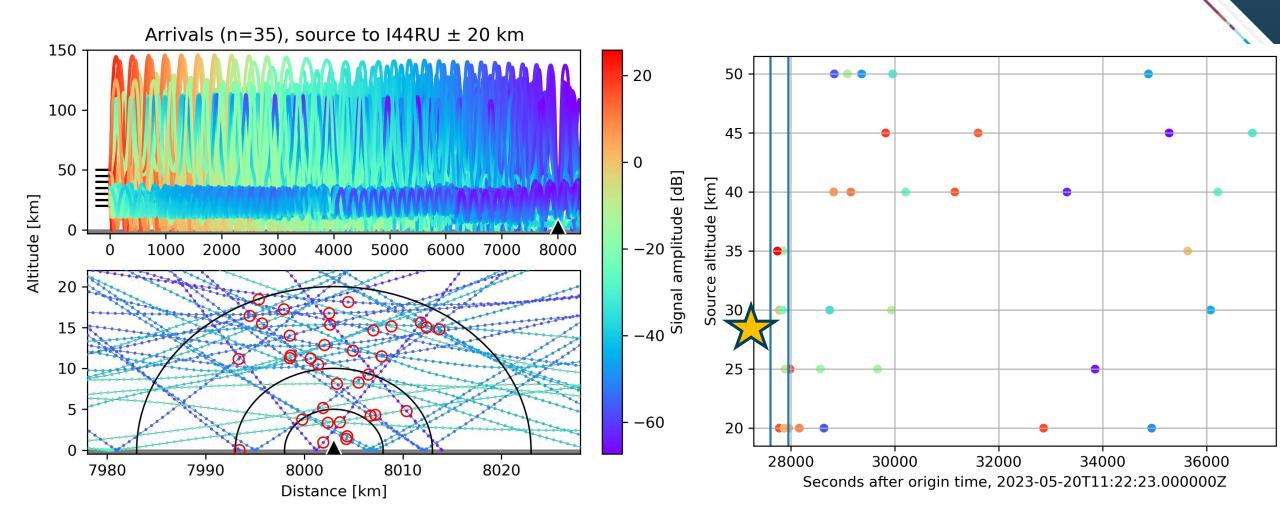


PROPAGATION MODELING



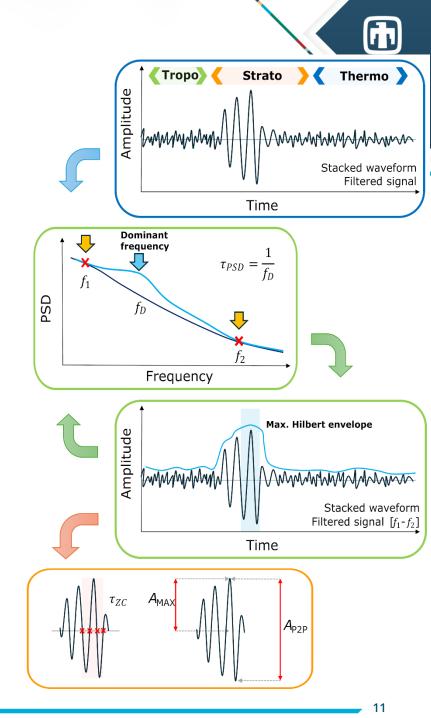
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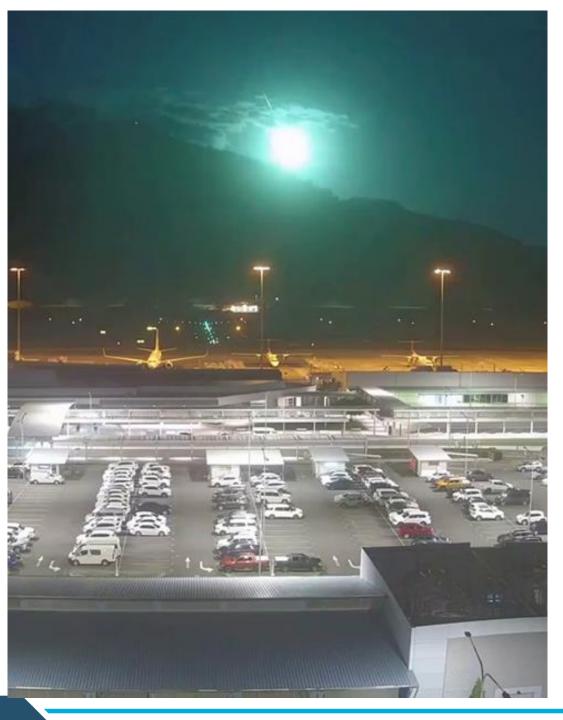
PROPAGATION MODELING



ENERGY ESTIMATE

- CNEOS derived energy from USG data: 7.2 kt of TNT equivalent
- Dominant signal period measured across infrasound stations can be utilized for estimating yield of a bolide
- We used the empirical energy relation from ReVelle (1997): log (E/2) = 3.34 log(T) - 2.58
- Fragmentation / airburst vs. cylindrical line source (shock geometry and size of blast radius are significantly different)
- Energy estimate from average signal period
 - 3.5 7.5 kt of TNT (ReVelle 1997)
 - 4.6 10.9 kt of TNT (Ens et al. 2012)
 - 13 30.3 kt of TNT (Gi & Brown 2017)
- Energy derived from various empirical relations can exhibit large uncertainties. In this case, infrasound-derived energy is in good agreement with the USG sensor-derived energy





SUMMARY

- The Australian bolide was detected from space and ground
- One of the most energetic events detected by USG sensors and the most energetic over Australia
- Different observational methods and theoretical approaches can complement each other
- Multi-modal sensing enables better characterization, and cross-validation
- Good agreement between energy estimate derived through infrasound detections and USG sensors

Stay tuned for a publication

ACKNOWLEDGEMENTS

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EXTRA SLIDES

