

### Solution of the "NPE-2019" by Using Inverse Atmospheric Transport Modelling

Sayed A. Mekhaimer

National Data Center for Monitoring Nuclear Tests (Egypt) National Research Institute of Astronomy and Geophysics (NRIAG)

CTBT: Science and Technology Conference 2023 - SnT2023

CTBT: Science and Technology Conference 2023

# Senario of NPE-2019

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ シのへで

CTBT: Science and Technology Conference 2023

### TRIGA reactor Accident on 30 July 2019

The national nuclear safety authority of the state of **RAETIA** released the following public announcement.

"An accident at TRIGA reactor facility located in Pavia, RAETIA, has occurred this morning 30th July 2019. We are expecting some small release of radioactive isotopes, but well below the hazardous limit for human health. A dedicated monitoring system has been activated around the facility and in the neighbouring in order to monitor the radioactivity in the air".



There is no need to activate any emergency procedures for the population neither any closure of schools and public areas is required"

#### Table: CS-134

#### Table: CS-137

Freq.

?

1447

6

< = > < = > < = > < = >

|        |                  |       | -         | S     | ΤN      | Collection Stop  | Conc. |
|--------|------------------|-------|-----------|-------|---------|------------------|-------|
| SIN    | Collection Stop  | Conc. | Freq.     | VI    | 200     | 2019-08-02 09:00 | 500   |
| VIP00  | 2019-08-02 09:00 | 2100  | ?         |       |         | 2019-08-06 23:15 | 6.2   |
|        | 2019-08-06 23:15 | 24.7  |           |       |         | 2019-08-07 23:29 | 4.8   |
|        | 2019-08-07 23:29 | 19.2  |           | LYI   | P41     | 2019-08-08 23:49 | 4.4   |
|        | 2019-08-08 23:49 | 17.4  |           |       |         | 2019-08-09 23:45 | 4.0   |
| L11 41 | 2019-08-09 23:45 | 15.8  | 2         |       |         | 2019-08-10 23:00 | 4.7   |
|        | 2019-08-10 23:00 | 18.7  | :         |       |         | 2019-08-09 06:00 | 3.4   |
|        | 2019-08-11 23:40 | 4.7   |           | KW    | P40     | 2019-08-10 06:00 | 4.8   |
|        | 2019-08-06 06:00 | 7.5   |           |       |         | 2019-08-11 06:00 | 4.4   |
|        | 2019-08-07 06:00 | 10.2  |           |       | ON DO I | 2019-08-10 23:29 | 7.4   |
|        | 2019-08-08 06:00 | 8.9   | Eukuchima | GNP21 | P21     | 2019-08-11 23:50 | 9.9   |
| NVF40  | 2019-08-09 06:00 | 13.5  | Fukushina |       |         |                  |       |
|        | 2019-08-10 06:00 | 19.1  |           |       |         |                  |       |
|        | 2019-08-11 06:00 | 17.4  |           |       |         |                  |       |
| NEP48  | 2019-08-09 09:00 | 5.0   | 0         |       |         |                  |       |
| RNP36  | 2019-08-10 00:00 | 3.5   | 0         |       |         |                  |       |
|        | 2019-08-10 00:00 | 7.1   |           |       |         |                  |       |
| NUF 33 | 2019-08-10 23:55 | 5.7   | 0         |       |         |                  |       |
| CNID01 | 2019-08-10 23:29 | 29.1  | 0         |       |         |                  |       |
| CNP21  | 2019-08-11 23:50 | 39.3  | U         |       |         |                  |       |
|        |                  |       |           |       |         |                  |       |

#### Table: CS-134

#### Table: CS-137

| STN    | Collection Stop  | Conc. | Freq.     |
|--------|------------------|-------|-----------|
| VIP00  | 2019-08-02 09:00 | 2100  | ?         |
|        | 2019-08-06 23:15 | 24.7  |           |
|        | 2019-08-07 23:29 | 19.2  |           |
|        | 2019-08-08 23:49 | 17.4  |           |
| LTF41  | 2019-08-09 23:45 | 15.8  | 2         |
|        | 2019-08-10 23:00 | 18.7  | f         |
|        | 2019-08-11 23:40 | 4.7   |           |
|        | 2019-08-06 06:00 | 7.5   |           |
|        | 2019-08-07 06:00 | 10.2  |           |
|        | 2019-08-08 06:00 | 8.9   | Fukuchimo |
| KWF40  | 2019-08-09 06:00 | 13.5  | Fukushima |
|        | 2019-08-10 06:00 | 19.1  |           |
|        | 2019-08-11 06:00 | 17.4  |           |
| NEP48  | 2019-08-09 09:00 | 5.0   | 0         |
| RNP36  | 2019-08-10 00:00 | 3.5   | 0         |
|        | 2019-08-10 00:00 | 7.1   | 0         |
| NUF 33 | 2019-08-10 23:55 | 5.7   | 0         |
| CNID01 | 2019-08-10 23:29 | 29.1  | 0         |
| GNP21  | 2019-08-11 23:50 | 39.3  | U         |

| STN    | Collection Stop  | Conc. | Freq. |
|--------|------------------|-------|-------|
| VIP00  | 2019-08-02 09:00 | 500   | ?     |
|        | 2019-08-06 23:15 | 6.2   |       |
|        | 2019-08-07 23:29 | 4.8   |       |
| LYP41  | 2019-08-08 23:49 | 4.4   | ?     |
|        | 2019-08-09 23:45 | 4.0   |       |
|        | 2019-08-10 23:00 | 4.7   |       |
|        | 2019-08-09 06:00 | 3.4   |       |
| KWP40  | 2019-08-10 06:00 | 4.8   | 1447  |
|        | 2019-08-11 06:00 | 4.4   |       |
| CNID21 | 2019-08-10 23:29 | 7.4   | 6     |
| GINFZI | 2019-08-11 23:50 | 9.9   | 0     |

The highest Concentration at VIP00. This may lead to assume a source near VIP00.

◆□ → ◆圖 → ◆臣 → ◆臣 → ○臣

#### Table: CS-134

#### Table: CS-137

| STN    | Collection Stop  | Conc. | Freq.     |
|--------|------------------|-------|-----------|
| VIP00  | 2019-08-02 09:00 | 2100  | ?         |
|        | 2019-08-06 23:15 | 24.7  |           |
|        | 2019-08-07 23:29 | 19.2  |           |
|        | 2019-08-08 23:49 | 17.4  |           |
| LIF4I  | 2019-08-09 23:45 | 15.8  | 2         |
|        | 2019-08-10 23:00 | 18.7  | f         |
|        | 2019-08-11 23:40 | 4.7   |           |
|        | 2019-08-06 06:00 | 7.5   |           |
|        | 2019-08-07 06:00 | 10.2  |           |
|        | 2019-08-08 06:00 | 8.9   | Eukushima |
| KWF40  | 2019-08-09 06:00 | 13.5  | Fukushima |
|        | 2019-08-10 06:00 | 19.1  |           |
|        | 2019-08-11 06:00 | 17.4  |           |
| NEP48  | 2019-08-09 09:00 | 5.0   | 0         |
| RNP36  | 2019-08-10 00:00 | 3.5   | 0         |
| DUDEE  | 2019-08-10 00:00 | 7.1   | 0         |
| nuP55  | 2019-08-10 23:55 | 5.7   | U         |
| CNID01 | 2019-08-10 23:29 | 29.1  | 0         |
| GNP21  | 2019-08-11 23:50 | 39.3  | U         |

| STN    | Collection Stop  | Conc. | Freq. |
|--------|------------------|-------|-------|
| VIP00  | 2019-08-02 09:00 | 500   | ?     |
|        | 2019-08-06 23:15 | 6.2   |       |
|        | 2019-08-07 23:29 | 4.8   |       |
| LYP41  | 2019-08-08 23:49 | 4.4   | ?     |
|        | 2019-08-09 23:45 | 4.0   |       |
|        | 2019-08-10 23:00 | 4.7   |       |
|        | 2019-08-09 06:00 | 3.4   |       |
| KWP40  | 2019-08-10 06:00 | 4.8   | 1447  |
|        | 2019-08-11 06:00 | 4.4   |       |
| CNID21 | 2019-08-10 23:29 | 7.4   | 6     |
| GNFZT  | 2019-08-11 23:50 | 9.9   | 0     |

The highest Concentration at VIP00. This may lead to assume a source near VIP00.

The second highest concentration detections at CNP21!

(日)

#### Table: CS-134

#### Table: CS-137

| STN    | Collection Stop  | Conc. | Freq.     |
|--------|------------------|-------|-----------|
| VIP00  | 2019-08-02 09:00 | 2100  | ?         |
|        | 2019-08-06 23:15 | 24.7  |           |
|        | 2019-08-07 23:29 | 19.2  |           |
|        | 2019-08-08 23:49 | 17.4  |           |
| LIF4I  | 2019-08-09 23:45 | 15.8  | 2         |
|        | 2019-08-10 23:00 | 18.7  | f         |
|        | 2019-08-11 23:40 | 4.7   |           |
|        | 2019-08-06 06:00 | 7.5   |           |
|        | 2019-08-07 06:00 | 10.2  |           |
|        | 2019-08-08 06:00 | 8.9   | Eukushima |
|        | 2019-08-09 06:00 | 13.5  | Fukushima |
|        | 2019-08-10 06:00 | 19.1  |           |
|        | 2019-08-11 06:00 | 17.4  |           |
| NEP48  | 2019-08-09 09:00 | 5.0   | 0         |
| RNP36  | 2019-08-10 00:00 | 3.5   | 0         |
| DUDEE  | 2019-08-10 00:00 | 7.1   | 0         |
| nuP33  | 2019-08-10 23:55 | 5.7   | U         |
| CNID01 | 2019-08-10 23:29 | 29.1  | 0         |
| UNP21  | 2019-08-11 23:50 | 39.3  | U         |

| STN    | Collection Stop  | Conc. | Freq. |
|--------|------------------|-------|-------|
| VIP00  | 2019-08-02 09:00 | 500   | ?     |
|        | 2019-08-06 23:15 | 6.2   |       |
|        | 2019-08-07 23:29 | 4.8   |       |
| LYP41  | 2019-08-08 23:49 | 4.4   | ?     |
|        | 2019-08-09 23:45 | 4.0   |       |
|        | 2019-08-10 23:00 | 4.7   |       |
|        | 2019-08-09 06:00 | 3.4   |       |
| KWP40  | 2019-08-10 06:00 | 4.8   | 1447  |
|        | 2019-08-11 06:00 | 4.4   |       |
| CNID21 | 2019-08-10 23:29 | 7.4   | 6     |
| GINFZI | 2019-08-11 23:50 | 9.9   | 0     |

- The highest Concentration at VIP00. This may lead to assume a source near VIP00.
- The second highest concentration detections at CNP21!
- There is a detection of Cs-134 on KWP40, 17 hours before LYP41?

(a)

#### Table: CS-134

#### Table: CS-137

| STN    | Collection Stop  | Conc. | Freq.     |
|--------|------------------|-------|-----------|
| VIP00  | 2019-08-02 09:00 | 2100  | ?         |
|        | 2019-08-06 23:15 | 24.7  |           |
|        | 2019-08-07 23:29 | 19.2  |           |
|        | 2019-08-08 23:49 | 17.4  |           |
| LIF4I  | 2019-08-09 23:45 | 15.8  | 2         |
|        | 2019-08-10 23:00 | 18.7  | f         |
|        | 2019-08-11 23:40 | 4.7   |           |
|        | 2019-08-06 06:00 | 7.5   |           |
|        | 2019-08-07 06:00 | 10.2  |           |
|        | 2019-08-08 06:00 | 8.9   | Eukushima |
| KWF40  | 2019-08-09 06:00 | 13.5  | Fukushima |
|        | 2019-08-10 06:00 | 19.1  |           |
|        | 2019-08-11 06:00 | 17.4  |           |
| NEP48  | 2019-08-09 09:00 | 5.0   | 0         |
| RNP36  | 2019-08-10 00:00 | 3.5   | 0         |
| DUDEE  | 2019-08-10 00:00 | 7.1   | 0         |
| NUF35  | 2019-08-10 23:55 | 5.7   | U         |
| CNID01 | 2019-08-10 23:29 | 29.1  | 0         |
| GNF21  | 2019-08-11 23:50 | 39.3  | U         |

| STN    | Collection Stop  | Conc. | Freq. |
|--------|------------------|-------|-------|
| VIP00  | 2019-08-02 09:00 | 500   | ?     |
|        | 2019-08-06 23:15 | 6.2   |       |
|        | 2019-08-07 23:29 | 4.8   |       |
| LYP41  | 2019-08-08 23:49 | 4.4   | ?     |
|        | 2019-08-09 23:45 | 4.0   |       |
|        | 2019-08-10 23:00 | 4.7   |       |
|        | 2019-08-09 06:00 | 3.4   |       |
| KWP40  | 2019-08-10 06:00 | 4.8   | 1447  |
|        | 2019-08-11 06:00 | 4.4   |       |
| CNID21 | 2019-08-10 23:29 | 7.4   | 6     |
| UNF 21 | 2019-08-11 23:50 | 9.9   | 0     |

- The highest Concentration at VIP00. This may lead to assume a source near VIP00.
- The second highest concentration detections at CNP21!
- There is a detection of Cs-134 on KWP40, 17 hours before LYP41?
- While Cs-137 is frequently measured in KWP40, but during 2-12 Aug. only three detections?

### Abnormal Radionuclide detections (LA-140 & BA-140)

LA-140 and BA-140 are never previously measured at stations: CNP21 and KWP40.

#### Table: LA-140

| STN    | Collection Stop Conc. |      |  |
|--------|-----------------------|------|--|
| VIP00  | 2019-08-02 09:00      | 1200 |  |
|        | 2019-08-06 23:15      | 11.1 |  |
|        | 2019-08-07 23:29      | 8.4  |  |
| LYP41  | 2019-08-08 23:49      | 7.1  |  |
|        | 2019-08-09 23:45      | 6.2  |  |
|        | 2019-08-10 23:00      | 6.9  |  |
|        | 2019-08-06 06:00      | 3.4  |  |
|        | 2019-08-07 06:00      | 4.4  |  |
| KWD10  | 2019-08-08 06:00      | 3.6  |  |
| NWF 40 | 2019-08-09 06:00      | 5.2  |  |
|        | 2019-08-10 06:00      | 6.9  |  |
|        | 2019-08-11 06:00      | 6.0  |  |
| CNID21 | 2019-08-10 23:29      | 10.5 |  |
| UNFZI  | 2019-08-11 23:50      | 13.7 |  |

#### Table: BA-140

| STN    | Collection Stop Conc. |      |  |
|--------|-----------------------|------|--|
| VIP00  | 2019-08-02 09:00      | 1000 |  |
|        | 2019-08-06 23:15      | 10.2 |  |
|        | 2019-08-07 23:29      | 7.7  |  |
| LYP41  | 2019-08-08 23:49      | 6.6  |  |
|        | 2019-08-09 23:45      | 5.7  |  |
|        | 2019-08-10 23:00      | 6.4  |  |
|        | 2019-08-06 06:00      | 3.1  |  |
|        | 2019-08-07 06:00      | 4.0  |  |
| KWD10  | 2019-08-08 06:00      | 3.3  |  |
|        | 2019-08-09 06:00      | 4.8  |  |
|        | 2019-08-10 06:00      | 6.4  |  |
|        | 2019-08-11 06:00      | 5.5  |  |
| CNID21 | 2019-08-10 23:50      | 9.7  |  |
| UNF21  | 2019-08-11 23:50      | 12.6 |  |

### Abnormal Radionuclide detections (LA-140 & BA-140)

LA-140 and BA-140 are never previously measured at stations: CNP21 and KWP40.

#### Table: LA-140

STN Collection Stop Conc. VIP00 2019-08-02 09:00 1200 2019-08-06 23 15 11.1 2019-08-07 23:29 8.4 LYP41 2019-08-08 23:49 7.1 2019-08-09 23:45 6.2 2019-08-10 23:00 6.9 2019-08-06 06:00 34 2019-08-07 06:00 4.4 2019-08-08 06:00 3.6 KWP40 5.2 2019-08-09 06:00 2019-08-10 06:00 6.9 2019-08-11 06:00 6.0 2019-08-10 23:29 10.5 CNP21 2019-08-11 23:50 13.7

#### Table: BA-140

| STN    | Collection Stop  | Conc. |  |
|--------|------------------|-------|--|
| VIP00  | 2019-08-02 09:00 | 1000  |  |
|        | 2019-08-06 23:15 | 10.2  |  |
|        | 2019-08-07 23:29 | 7.7   |  |
| LYP41  | 2019-08-08 23:49 | 6.6   |  |
|        | 2019-08-09 23:45 | 5.7   |  |
|        | 2019-08-10 23:00 | 6.4   |  |
|        | 2019-08-06 06:00 | 3.1   |  |
|        | 2019-08-07 06:00 | 4.0   |  |
|        | 2019-08-08 06:00 | 3.3   |  |
|        | 2019-08-09 06:00 | 4.8   |  |
|        | 2019-08-10 06:00 | 6.4   |  |
|        | 2019-08-11 06:00 | 5.5   |  |
| CNID21 | 2019-08-10 23:50 | 9.7   |  |
| GNEZI  | 2019-08-11 23:50 | 12.6  |  |

The highest Concentration at VIP00. The second highest concentration detections at CNP21.

### Abnormal Radionuclide detections (LA-140 & BA-140)

LA-140 and BA-140 are never previously measured at stations: CNP21 and KWP40.

#### Table: LA-140

STN Collection Stop Conc. VIP00 2019-08-02 09:00 1200 2019-08-06 23 15 11.1 2019-08-07 23:29 8.4 LYP41 2019-08-08 23:49 7.1 2019-08-09 23:45 6.2 2019-08-10 23:00 6.9 2019-08-06 06:00 34 2019-08-07 06:00 4.4 2019-08-08 06:00 3.6 KWP40 2019-08-09 06:00 5.2 2019-08-10 06:00 6.9 2019-08-11 06:00 6.0 2019-08-10 23:29 10.5 CNP21 2019-08-11 23:50 13.7

#### Table: BA-140

| STN    | Collection Stop  | Conc. |  |
|--------|------------------|-------|--|
| VIP00  | 2019-08-02 09:00 | 1000  |  |
| LYP41  | 2019-08-06 23:15 | 10.2  |  |
|        | 2019-08-07 23:29 | 7.7   |  |
|        | 2019-08-08 23:49 | 6.6   |  |
|        | 2019-08-09 23:45 | 5.7   |  |
|        | 2019-08-10 23:00 | 6.4   |  |
|        | 2019-08-06 06:00 | 3.1   |  |
|        | 2019-08-07 06:00 | 4.0   |  |
| KWP40  | 2019-08-08 06:00 | 3.3   |  |
|        | 2019-08-09 06:00 | 4.8   |  |
|        | 2019-08-10 06:00 | 6.4   |  |
|        | 2019-08-11 06:00 | 5.5   |  |
| CNID21 | 2019-08-10 23:50 | 9.7   |  |
| GINF21 | 2019-08-11 23:50 | 12.6  |  |

The highest Concentration at VIP00. The second highest concentration detections at CNP21.

There are detections of both LA-140 and BA-140 on KWP40, 17 hours before LYP41?

### **Xenon Detections**

#### Table: Xenon Istopes Detections During NPE-2019

| STN   | Collection Stop | Xe-133 | Xe-135 | Xe-131m | Xe-133m |
|-------|-----------------|--------|--------|---------|---------|
|       | 2019/08/02 06   | 22.6   | 6.39   |         | 1.53    |
|       | 2019/08/03 06   | 9.6    | 1.1    |         |         |
| DEX33 | 2019/08/04 06   | 16.6   |        |         |         |
|       | 2019/08/05 06   | 4.9    |        |         |         |
|       | 2019/08/06 06   | 0.97   |        |         |         |
| NEX48 | 2019/08/09 06   | 0.44   |        |         |         |
|       | 2019/08/02 06   | 52     | 15     | 1.5     | 3.5     |
|       | 2019/08/03 06   | 121    | 7      |         | 6.6     |
| VIX00 | 2019/08/04 06   | 24     |        |         |         |
|       | 2019/08/05 06   | 6.4    |        |         |         |
|       | 2019/08/06 06   | 6.7    |        |         |         |
|       | 2019/08/07 06   | 93     |        | 1.6     | 2.      |
|       | 2019/08/08 06   | 22     |        |         |         |

#### Table: Freq. of Xenon detections

| STN   | lso.    | Freq. |
|-------|---------|-------|
| DEX33 | Xe-133  | 2909  |
| DEX33 | Xe-135  | 30    |
| DEX33 | Xe-131m | 6     |
| DEX33 | Xe-133m | 18    |

▲□▶ ▲□▶ ▲□▶ ▲□▶ ▲□ シのへで

# Atmospheric Transport Modeling

CTBT: Science and Technology Conference 2023















### Adjoint Atmospheric Transport Modeling From IMS



CTBT: Science and Technology Conference 2023



### Adjoint Atmospheric Transport Modeling From IMS



CTBT: Science and Technology Conference 2023









### Inverse Modelling Methodology

The adjoint transport and dispersion equation has the following form in the Lagrangian specification:

$$-\frac{D\overline{C^*}}{Dt} - \nabla \cdot (K\nabla \overline{C^*}) = R \tag{1}$$

Where the observed concentration values are related to the *source-receptor-sensitivity* (SRS) field as follows

$$C_k^{obs} = \sum_n \sum_j \sum_j S_{ijn} C_{ijnk}^* \quad k=1,2,\cdots,K$$
(2)

where K is the number of samples and  $S_{ijn}$  is the spatial-temporal source field. The estimation of modeled concentrations values

$$\boldsymbol{C}_{ijnk}^{est} = \alpha_{ijn} \boldsymbol{C}_{ijnk}^* + \beta_{ijn} \tag{3}$$

### Estimation of Likelihood

$$\mathbb{P}(C|S_{ijn}(s), I) \propto exp\left(-\frac{1}{2}\sum_{k}\frac{(C_{k}^{obs} - C_{ijnk}^{est}(s))^{2}}{(\sigma_{k}^{obs})^{2} + (\sigma_{k}^{est})^{2}}\right)$$
(4)

イロト イポト イヨト イヨト 三日

where the likelihood function,  $\mathbb{P}(C|S_{ijn}(s), I)$ , represents the probability that the source,  $S_{ijn}(s)$ , which is located at grid point *i*, *j* and released during time interval *n* with fixed rate *s* will give estimated concentrations  $C_{ijnk}^{est}(s)$ , which matches the observed data  $C_k^{obs}$ .

### **Bayesian Approach**

$$P(S|C, I) = \frac{P(S|I)\mathbb{P}(C|S, I)}{P(C|I)}$$

I represents the prior information,

- C is concentration measurements,
- S is the source parameters,
- P(S|I) is the prior PDF of the source, which is the conditional PDF of source parameters S, given only the prior information I.
- P(C|S, I) is the likelihood function, which measures how good the fitting between the observed and the modeled concentration, given the assumed source parameter S and the prior information.
- P(S|C, I) is the posterior PDF of the source parameters, given the measurements C and the prior information.
- P(C|I) is the marginal distribution, which is obtained by summation the numerator of Eq. 5 overall assumed values of S. Thus, it is independent of S, and it can be considered as a normalization constant.

$$P(C|I) = \sum_{S} P(S|I) \mathbb{P}(C|S, I)$$
(6)

(5)

# Results

<□> 
<□> 
<□> 
E> 
E

CTBT: Science and Technology Conference 2023

















































































































































# **SHI** Analysis

#### By Mona Abdel Azem, Adel Samy, Islam Hamama

(日)

CTBT: Science and Technology Conference 2023

#### Seismic and Infrasound Detections

To search for possible events related to the unusual radionuclide detections, the IMS stations and Eastria National SHI data are used and we found the following event:

#### Seismic Event:

| Date_time         | Latitude     | Longitude | Magnitude | Depth |
|-------------------|--------------|-----------|-----------|-------|
| 2019/07/29 23:17: | 47.3 47.8660 | 9.2899    | MI 4.4    | 0     |



• • • • • • • • • • • •

#### Seismic and Infrasound Detections

To search for possible events related to the unusual radionuclide detections, the IMS stations and Eastria National SHI data are used and we found the following event:

#### • Two Infrasound Events:

| Date_time           | Latitude | Longitude | Date time             | Latitude | Longitude |
|---------------------|----------|-----------|-----------------------|----------|-----------|
| 2019/07/30 09:07:31 | 44.2204  | 16.39     | 2019/08/01 08:00:37.5 | 44.3006  | 16.5513   |



#### Summary of SHI detections



◆□ → ◆圖 → ◆臣 → ◆臣 → ○臣
# Source Determination

▲□▶ ▲□▶ ▲三▶ ▲三▶ 三三 - つくで

# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)


# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)



# Relative Likelihood using Flexpart-ECMWF (Ba-140)


#### Relative Likelihood using Flexpart-ECMWF (Ba-140)



#### Relative Likelihood using Flexpart-ECMWF (Ba-140)



#### Possible Source Regions (using La-140 & Ba-140)



LA-140 Sum (12 UTC 29-30 July 2019) ecmwf









#### Possible Source Regions (using Cs-134 & Cs-137)











### Ensemble Mean Likelihood (All Nuclides & Modoles)

Under assumption that all nuclides are release from the same sources, we can calculate the ensemble mean using all nuclides and models.



## Ensemble Mean Likelihood (All Nuclides & Modoles)

Under assumption that all nuclides are release from the same sources, we can calculate the ensemble mean using all nuclides and models.



Sum Likelihood (12 UTC 29 July and 12 UTC 30 July 2019)

CTBT: Science and Technology Conference 2023

# Marginal Temporal PDF



イロト イポト イヨト イヨト

э

Posterior Probability

# Marginal Longitudinal PDF



Posterior Probability

## Marginal Meridional PDF



Image: A math a math

Posterior Probability

# **Isotopic Ratios**

・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・
・

## Cs-134/Cs-137 Activity Ratio



э

#### Cs-134/Cs-137 Activity Ratio



 The Cs-134/Cs-137 Activity Ratio for stations:LYP41, KWP40, and CNP21 are between 3.9506 and 3.9777, while that for VIP00 is 4.205.

• • • • • • • • • • • •

#### Cs-134/Cs-137 Activity Ratio



- The Cs-134/Cs-137 Activity Ratio for stations:LYP41, KWP40, and CNP21 are between 3.9506 and 3.9777, while that for VIP00 is 4.205.
- The ratios for those stations except VIP00 are fitting the Cs-134/Cs-137 decay curve with initial yield ratio 0.2735.

#### Cs-134/Cs-137 Activity Ratio



- The Cs-134/Cs-137 Activity Ratio for stations:LYP41, KWP40, and CNP21 are between 3.9506 and 3.9777, while that for VIP00 is 4.205.
- The ratios for those stations except VIP00 are fitting the Cs-134/Cs-137 decay curve with initial yield ratio 0.2735.
- The analysis of Cs-134/Cs-137 Activity Ratio suggests that the detections for VIP00 are from different source than that of other stations.

#### La-140/Ba-140 Activity Ratio



3

#### La-140/Ba-140 Activity Ratio



 La-140/Ba-140 Ratio for stations:LYP41, KWP40, and CNP21 are between 0.982 and 0.987, while that for VIP00 is 1.053.

Image: A matching of the second se

- E - N

### La-140/Ba-140 Activity Ratio



- La-140/Ba-140 Ratio for stations:LYP41, KWP40, and CNP21 are between 0.982 and 0.987, while that for VIP00 is 1.053.
- The difference of La-140/Ba-140 ratio for VIP00 from other stations suggests that its detections are from different source than the source causes the detections in LYP41, KWP40, and CNP21.
- By assuming that source related to VIP00 detections is nuclear test, the zero release time will 6.83 days (27 July).

# Conclusion

◆□→ ◆御→ ◆恵→ ◆恵→ □ 恵

 The particulate detections in the NPE-2019 exercise represent an unusual nuclear event not previously seen in IMS history except Fukushima 2011.

(日) (圖) (E) (E) (E)

- The particulate detections in the NPE-2019 exercise represent an unusual nuclear event not previously seen in IMS history except Fukushima 2011.
- The xenon detections in the exercise are significant. However, quite similar detections previously occurred in this geographical region.

< ロ > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

- The particulate detections in the NPE-2019 exercise represent an unusual nuclear event not previously seen in IMS history except Fukushima 2011.
- The xenon detections in the exercise are significant. However, quite similar detections previously occurred in this geographical region.
- The inverse modelling by using all anomaly particulates detections shows that there is another possible location north of TRIGA reactor as another source.

(日) (個) (E) (E) (E)

- The particulate detections in the NPE-2019 exercise represent an unusual nuclear event not previously seen in IMS history except Fukushima 2011.
- The xenon detections in the exercise are significant. However, quite similar detections previously occurred in this geographical region.
- The inverse modelling by using all anomaly particulates detections shows that there is another possible location north of TRIGA reactor as another source.
- The radionuclide isotopic ratios analysis shows that there are two different sources at least.

(日) (個) (E) (E) (E)

- The particulate detections in the NPE-2019 exercise represent an unusual nuclear event not previously seen in IMS history except Fukushima 2011.
- The xenon detections in the exercise are significant. However, quite similar detections previously occurred in this geographical region.
- The inverse modelling by using all anomaly particulates detections shows that there is another possible location north of TRIGA reactor as another source.
- The radionuclide isotopic ratios analysis shows that there are two different sources at least.
- The Seismic and infrasound analysis techniques determine three shock waves in the area of interest. Two purely infrasound and one seismic event. Only the North pure Seismic location (47.7° N, 9.29° E) is consistent with the ATM inverse modeling.

• This location is consistent with atmospheric inverse modelling using all available particulate detections, where the release may occur during 29-30 July 2019.

(日) (圖) (目) (E) (E) (E)

- This location is consistent with atmospheric inverse modelling using all available particulate detections, where the release may occur during 29-30 July 2019.
- Therefore, We can conclude that there were two nuclear events:
  - The announced TRIGA reactor accident, and
  - unannounced release from location (47.7° N, 9.29° E), with release occurred after an explosion on 29 July at 23:17.

(日)

- This location is consistent with atmospheric inverse modelling using all available particulate detections, where the release may occur during 29-30 July 2019.
- Therefore, We can conclude that there were two nuclear events:
  - The announced TRIGA reactor accident, and
  - unannounced release from location (47.7° N, 9.29° E), with release occurred after an explosion on 29 July at 23:17.

• While the adopted approach assumes that one source, it can be used to determine multi-sources (two in the current case) according to the spatial distribution of the likelihood function.

# Thank You